## P. L. A I N

## TRIGONOMETRY

RENDERED EASY AND FAMILIAR,

BY CALCULATIONS IN ARITHMETICK ONLY:

WITH ITS

## APPLICATION AND USE

IN ASCERTAINING

All Kinds of Heights, Depths, and Distances,

IN

The MEAVENS, as well as on the EARTH and SEAS;

WHETHER OF

TOWERS, FORTS, TREES. PYRAMIDS, COLUMNS, WELLS, SHIPS, HILLS, CLOUDS, THUNDER AND LIGHTNING, ATMOSPHERE, SUN, MOON, PLANETS, MOUNTAINS IN THE MOON, SHADOWS OF EARTH AND MOON, BEGINNING AND END OF ECLIPSES, &c.

IN WHICH IS ALSO SHEWN,

A Curious Trigonometrical Metrop of discovering the Places where Bees hive in LARGE WOODS, in Order to obtain, more readily, the SALULARY PRODUCE of those LITTLE INSECTS.

## By the Rev. R. TURNER, of Magdalen Hall, Oxford;

Author of The View of the Earth, View of the Heavens; -System of Gauging; -and Chronologer Perpetral:

Rector of Comberton, Vicar of Elmley, Minister of Smoulton, and Chaplain to the Right Honourable the Countess Dowager of WIGTON.

Cuncla Trigonus habet, satagitquædocla Mathelis, Ille aperit clausum, quicquid Olympus habet.

Within the grand Triangle lies unweil'd, What Sages fought for, and what Heaven conreal'd.

#### LONDON:

PRINTED FOR S. CROWDER, THE PATER-NOSTER ROW. 1778.

# GENTLEMEN,

WHOSE GENIUS MAY INCLINE, OR EMPLOYMENT LEAD THEM TO THE STUDY OF THE

## MATHEMATICKS.

GENTLEMEN,

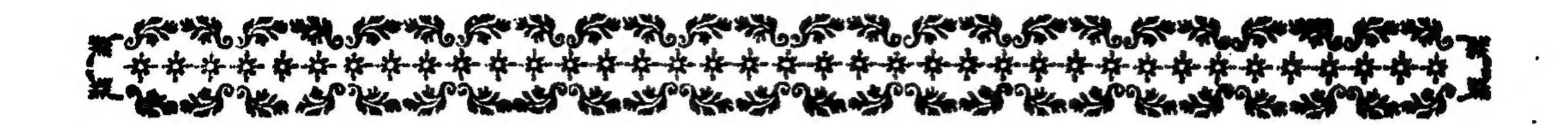
RIGONOMETRY has always been look'd on as one of the most useful Branches of Mathematical Learning. Navigation, Surveying, Astronomy, &c. stand wholly upon this Basis. But the common Method of answering these Problems by large Tables of Sines, Tangents, and Secants, renders it not only expensive by the Purchase of them; but often precarious in the Solution, through Mistakes of the Press. I have therefore, for the Use of the Young Mathematician, (from a Consideration of what has been published on this curious Subject) composed the present System, by which any of the Cases in Right or Oblique Plain Triangles may be answered on the Spot, by an easy Calculation in Arithmetick only \*. The great Advantages resulting from this Method to Gentlemen in the Army or Navy, as well as to Those in their private Studies at Home, must immediately appear; as it will be found to answer the most necessary Problems as expeditiously as Logarithms, (oftentimes more so;) and at the same Time wholly deliver you from those voluminous Tables and the inartificial Fatigues of carrying them always with you.—Should this little Treatise be so happy as to meet your Approbation, it will give a particular Pleasure to,

Your most humble Servant,



The AUTHOR.

It is wished that some Gentleman would undertake the Publication of Spherical Trigonometry in a similar Manner.

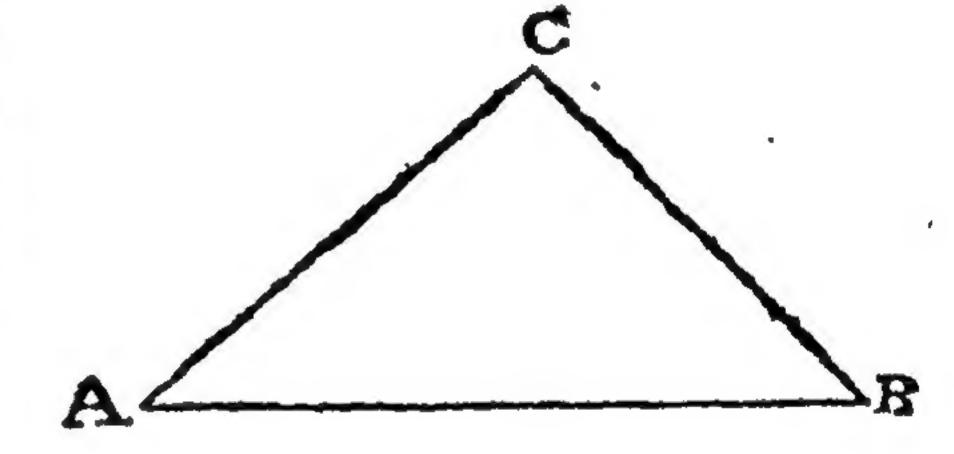


## P L A I

## TRIGONOMETRY.

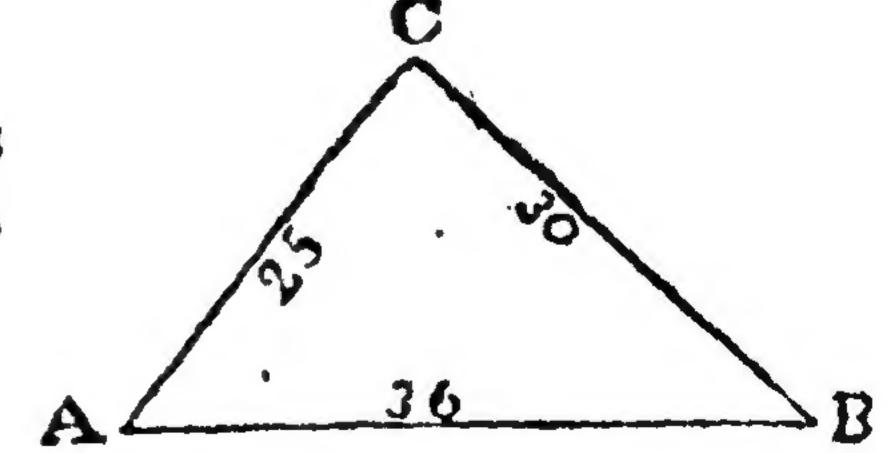
RIGONOMETRY is that Part of Mathematicks, which is employed in calculating the Sides and finding the Angles of any Triangle required; it is of the greatest Use, as nothing in Navigation, Astronomy, &c. can be done without it; and depends on the Knowledge of the following Observations, or Properties of that Figure.

(1st.) Every Triangle consists of Six Parts; that is, of *Three Sides* and *Three Angles*, as in the Figure ABC; the Three Sides are, AB, AC, CB, and the Three Angles, A, B, C.



Note. Sometimes an Angle is expressed by Three Letters; in that Case, the Middle Letter denotes the Angular Point. Thus, ABC expresses the Angle B; BAC the Angle A; and ACB the Angle C.

(2d.) The Sides of all plain Triangles are measured by a Line of equal Parts, as of Inches,—Feet,—Yards,—or Leagues.

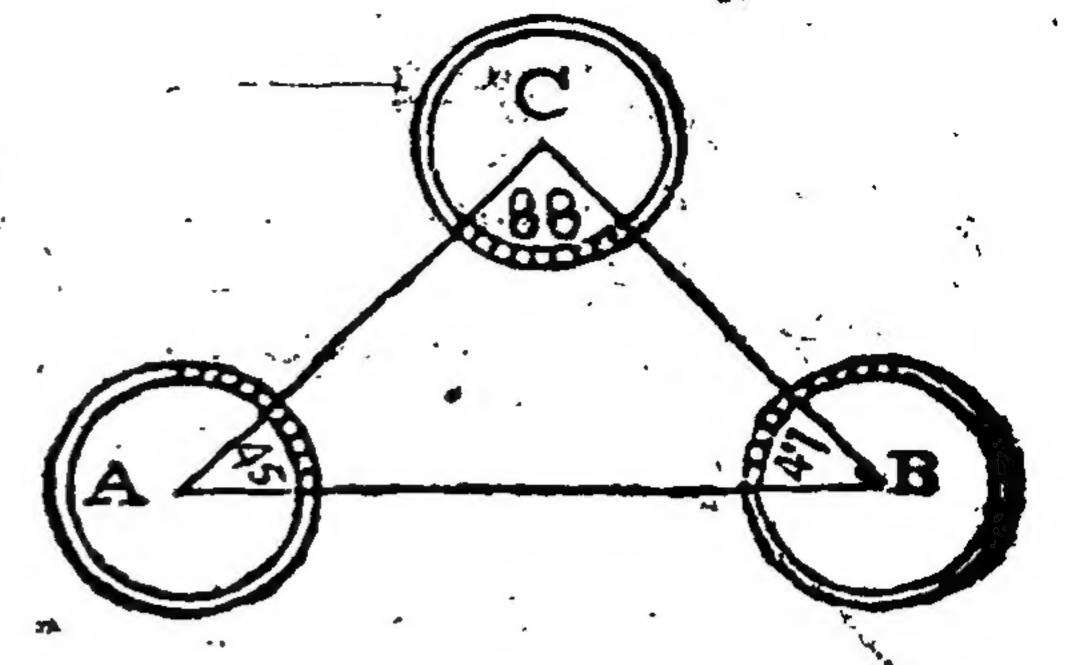


Thus, The Side AB is 36 Leagues.—The Side AC 25 Leagues.—And the Side BC is 30 Leagues.

(3d.), The

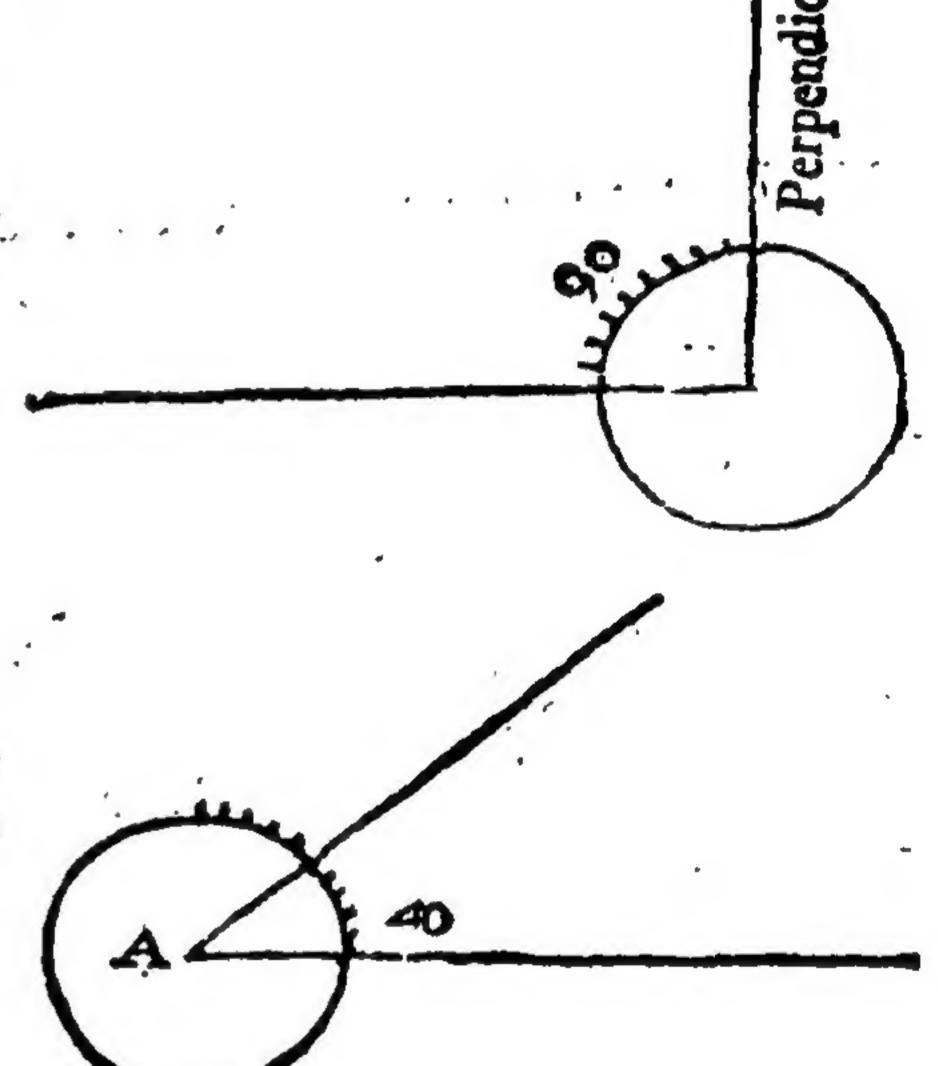
## PLAIN TRIGONOMETRY

(3d.) The Angles are measured by the Arch of a Circle described upon the Angular Point, and contained between the Two Legs that form the Angle.



Note. Every Circle is divided into 360 equal Parts, called Degrees; each of which is divided into 60 more, called Minutes: And the Number of Degrees contained between the Two Legs, that constitute the Angle, is the Measure of that Angle. Thus, The Angle A is 45 Degrees.—The Angle B 47.—The Angle C 88.

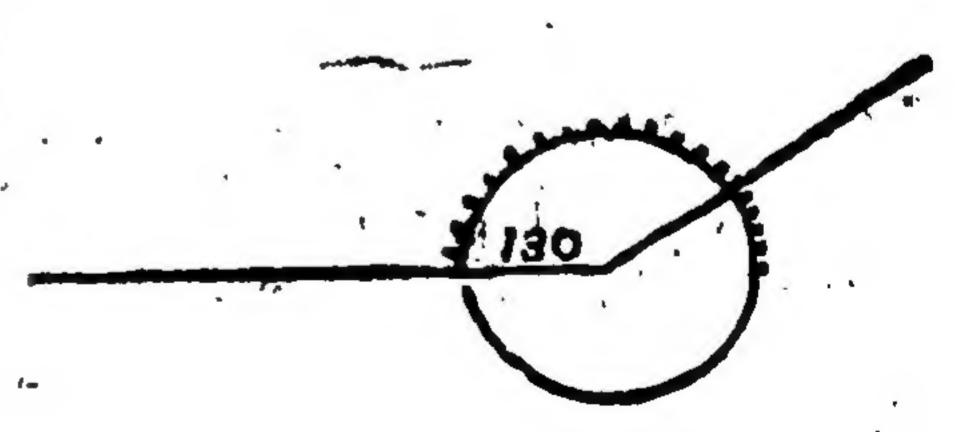
(4th.) If the Arch of a Circle intercepted between the Two Legs be exactly 90 Degrees, the Angle is called a Right Angle, and the Legs are perpendicular to one another.



(5th.) If the Arch of the Circle be less than 90 Degrees, the Angle is said to be Acute.

Note. What an Acute Angle wants of 90 Degrees, is called the Complement of that Angle. Thus, Suppose the Angle A was 40 Degrees; then its Complement is 50 Degrees; for 40 added to 50 make 90, as observed before.

(6th.) If the Arch of a Circle be more than 90 Degrees, the Angle is said to be Obtuse; and so continues to 180 Degrees, where the Angle vanishes, the Lines becoming Strait.

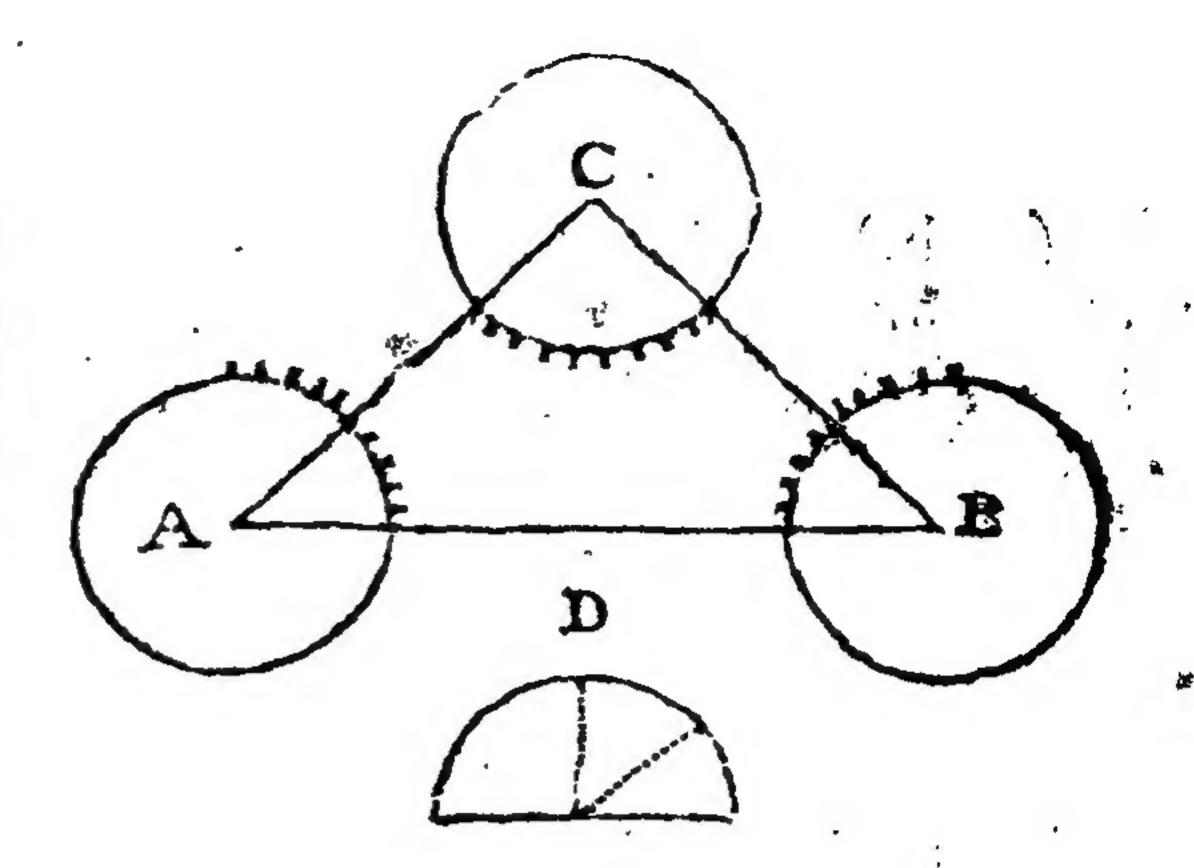


NOTE. What an Obtuse Angle wants of 180 Degrees, is also called the Complement of that Angle to a Semicircle.

(7th.) The

### RENDERED EASY AND FAMILIAR.

(7th.) The Three Angles of every plain Triangle, being taken together, make 180 Degrees (equal to a Semicircle), and this they always do, let the Triangle be drawn however you please.

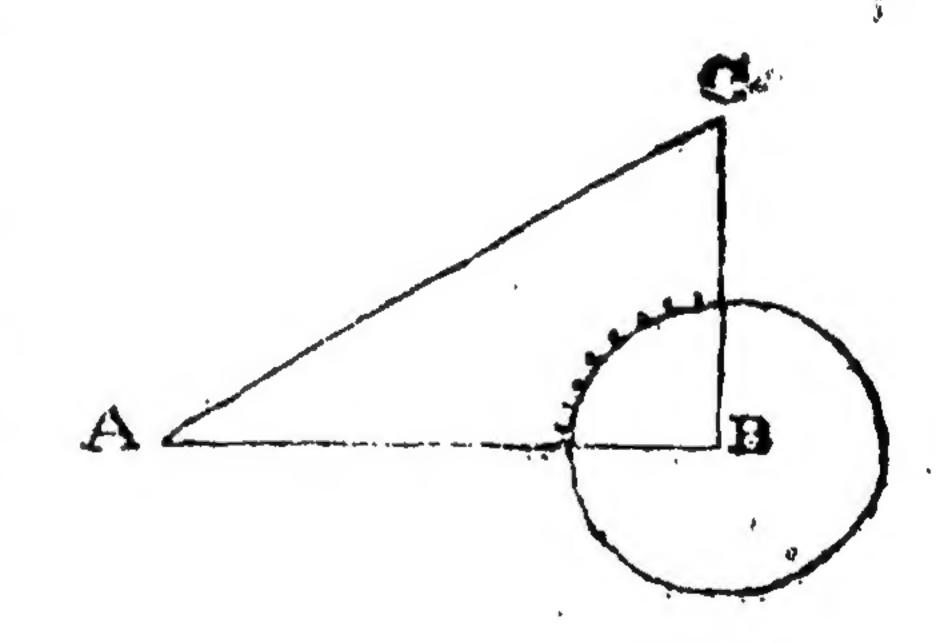


Thus, If the Semicircle D be drawn with the same Radius, or opening of the Dividers, as the little Circles on the Angles A, B, C, are; you will find, by taking off the several Arches, and applying them to the Semicircle, that they will just fill it up, and thereby make 180 Degrees; because every Semicircle contains that Number of Degrees.

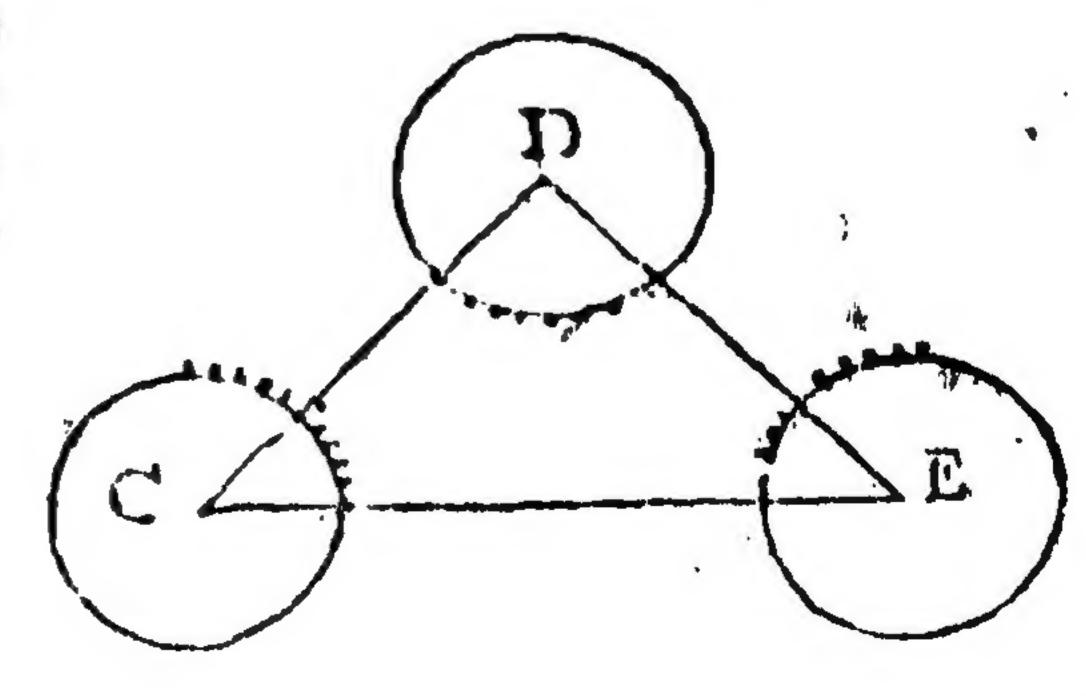
Hence it is evident, that if one Angle be a Right One, the other Two will be Acute; and taken together, be equal to one Right Angle, or just 90 Degrees.

Hence also, if Two Angles of any Triangle are known, the Third is easily found, being only the Degrees the other Two Angles want of 180.

(8th.) If a Triangle has one Right Angle, it is called a Right Angled Triangle: Thus ABC is a Right Angle Triangle, Right-angled at B.---In all Right Angle Triangles, the longest Leg is called the Hypothenuse; --- the Leg on which it stands, the Base; --- and the other Leg, the Perpendicular.



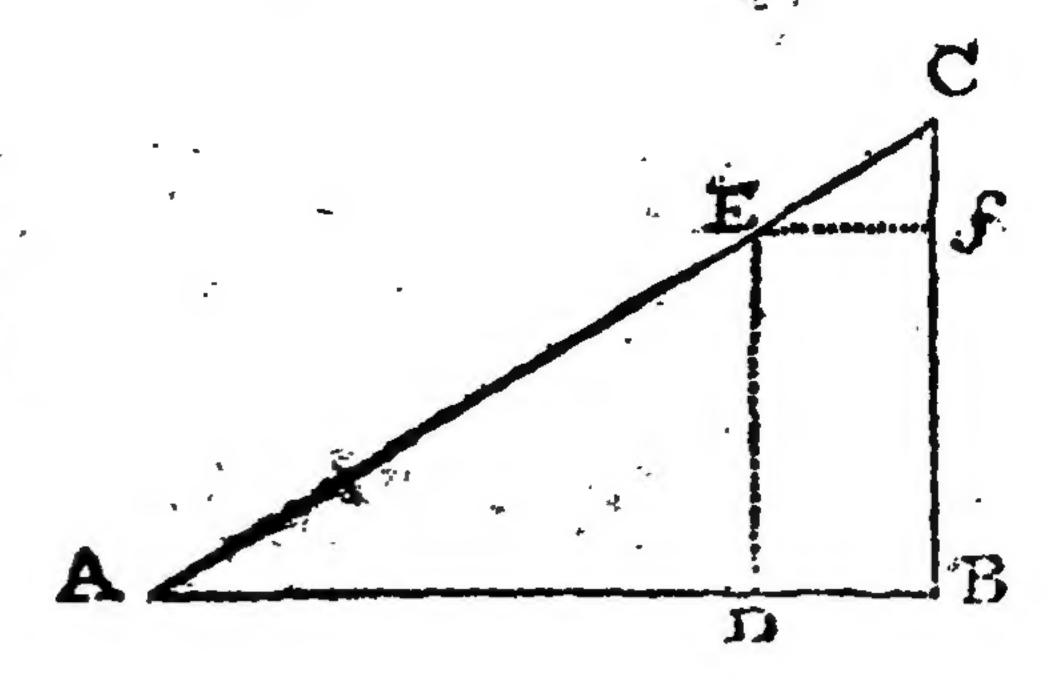
(9th.) If neither of the Angles is a Right One, then it is called an Oblique Angled Triangle; as the Triangle CDE is an Oblique Triangle.



(10th.) If

## PLAIN TRIGONOMETRY

(10th.) If the Angles of one Triangle are equal to the Angles of another Triangle, the Sides of the former are proportioned to the Sides of the atter.



Thus, If in the Triangle ABC, you draw the Line ED parallel to CB, the smaller Triangle ADE will be similar to, i. e. will have the same Angles with the larger Triangle ABC .-- It will therefore always hold, as

AB : AD : : BC : DE ... Or, as AB : BC : : AD : DE. Or, as AD: DE:: AB : AC. Or, as AD: DE:: Ef: fC, &c.

(11th.) In all Triangles the greatest Side is opposite to the greatest Angle; and on the contrary, the greatest Angle is opposite the greatest Side.---If Two Sides are equal, the opposite Angles are equal.----If all the Sides are equal, then all the Angles are equal to each other.

(12th.) In all Triangles, every Side is in proportion to its opposite Angle, and every Angle to its opposite Side: And further, as the Angle opposite to one Side, is to the Angle opposite the other Side, so are the Sides themselves one to another; and the contrary, the Sides to the Angles.

Every Triangle, as I observed before, consists of Six Parts---Three Sides and Three Angles. If any Three of the Six Parts (excepting the Three Angles) are given, any one, or every one of the rest may be found, without the painful Deductions and voluminous Tables of Logarithms, Sines, Tangents, and Secants, by the following Rules and Axioms ".

The foregoing Properties of a Triangle are so manifest, that they stand in need of no surther Illustra-

tion or Demenstration.

\*

This Method will be found as exact, as that by the Logarithms, if you carry on the Operation to Three or Feur Decimal Places; but for common Purposes, One or Two Decimal Places will be near enough. You must also remember to reduce the Minutes (and Seconds) of the Angles to Decimals of a Degree, which is eafily done, by allowing One Tenth for every Six Minutes. ---- Or you may turn the Minutes into Decimals thus: As 60, the Minutes in one Degree, : are to the Minutes given, : : so are 10, 100, 1000, &c. to the Decimal required.

## Of Right Angled TRIANGLES.

THERE are generally reckoned by Writers on this Subject Seven Cases; but by this Method they are all reduced to Four; the Solutions of which depend on the following Axioms.

AXIOM I. Divide 4 Times the Square of the Complement of the Angle, whose opposite Side is either given or sought, by 300 added to 3 Times the said Complement; this Quotient added to the said Angle, will give you an Artisicial Number, called sometimes the Natural Radius\*, which will ever bear the same Proportion to the Hypothenuse, as that Angle bears to its Side.——In Angles under 45 Degrees, the Artisicial Number may be sound thus: Divide 3 Times the Square of the Angle itself, whose opposite Side is given or sought, by 1000; the Quotient added to 57.3 †, a fixed Number, that Sum will be the Artisicial Number required.—This is to be used, when the Angles and a Side are given, to find another Side.

This Axiom is derived from collecting together a few leading Terms of a swift converging Series for determining the Length of the Sine or Co-sine, from the Length of a given Arch. Thus, If x represent the versed Sine, and z the Length of the correspondent Arch, the Radius being r, we shall have  $x = \frac{x^2}{2r} - \frac{x^4}{24r^3}$  (see Simpson's Fluxions, p. 500) and, consequently, the Cosine of the Arch  $z = r + \frac{z^4}{24r^3} - \frac{z^2}{2r}$ Now if we suppose r=1, the said Co-sine will be  $1+\frac{z^4}{24}-\frac{z^2}{2}$ which gives this rule: "Square the Complement of the given Arch. " and divide it by 2, subtract one Sixth of the Quantity squared from " the Quotient, and the Remainder from the Radius = 1. the Resi-"due will be the Sine of the proposed Arch." This Rule will give the Sine true to three Places in Decimals; but the Length of the Arch measuring the Angle must be taken in Terms of the Radius: In Order to remove this Difficulty let the Radius be increased to 57.3 nearly equal to the Radius of a Circle, whose Circumference is 360 equal Parts. Substitute this Value for r in the above Expression,  $r - x = r + \frac{x^4}{24^{1/4}} - \frac{x^2}{2r}$ , and we shall have  $57.3 + \frac{x^4}{24 \times 57.3} - \frac{x^2}{2 \times 57.3}$  for the Cosine of the Arch z, which may now be expressed by the Degrees and Decimal Parts, measuring the given Angle.

The Natural Radius is only turning the Right Angle, = 90 Degrees, into an artificial Number, which shall always bear the same Proportion to the Hypothenuse, as the given Angle does to its opposite Leg. † 57.3 is the Radius of a Circle whose Circumference is 360, or more exact, 57 25979.

From this Investigation the first Part of the above Axiom is easily deduced, as will appear by expressing that Rule algebraically, in Terms of the Radius 57.3, z being the Co-sine of the given Angle; for it will stand thus:  $\frac{4z^2}{3\infty+3z} + 90 - z$ , the Radius; consequently, it will be, as  $90 - z : \frac{4^{\frac{2^2}{300+3^2}}}{300+3^2} + 90 - z ::$  the Sine of the Angle expressed by 90 - z : 57.3, the Radius of a Circle whose Circumference is 360 equal Parts. Q. E. D.

The second Part of this Axiom may be easily investigated in the following Manner: Let z = the Length of any Arch, expressed in Terms of r, the Radius, and y = the Sine of that Arch; then will  $y = z - \frac{z^3}{2 \cdot 3r^2} + &c$ : And if we put r = 1, then will  $y = z - \frac{z^3}{2 \cdot 3}$ ; or, which is the same Thing  $y = 1 - \frac{z^2}{6} \times z$ , be the Sine of that Arch expressed in Terms of the Radius. (See Simpson's Fluxions, p. 501) But in Order to express it in Terms of the Angle itself, we must put r = 57.3, the Radius of a Circle whose Circumserence is 360 equal Parts. Then will the above Series, when the Angle is less than 45 Degrees, give an Expression nearly to  $\frac{3^{25}}{1000} + r = y$ , the Rule laid down in the second Part of this Axiom. Q. E. D.

AXIOM II. The Square of both the Legs, i. e. the Square of the Base and Perpendicular added together, is equal to the Square of the Hypothenuse; whose Root is the Hypothenuse itself.—This is made use of, when the Base and Perpendicular are given, to find the Hypothenuse.

This is the 47th Proposition of Euclid's 1st Book, and may be easily demonstrated in the following Manner.

Let ABC be a Right Angled Triangle, and let the Base AB = a, the Perpendicular BC = b, and the Hypothenuse AC = x; then will xx = aa + bb.

#### DEMONSTRATION.

Continue the Perpendicular BC till CD = AB, and on BD, as a Base, draw the Square BDFE. Upon AC, as a Base form another Square, as ACHG. Then will the Area of the great Square, BDFE == the Area of the four Triangles ABC, CDH, HFG, GAE + the Area of the Square ACHG. But the Area of each Triangle is  $\frac{ab}{2}$ ; consequently, the Area of the four Angles is 2 ab. Therefore, the Area of the great Square is xx + 2ab. But the Area of the great Square is also = the Rectangle of a + b, or aa + 2ba + bb; consequently, xx + 2ba = aa + 2ba + bb. Expunge 2ba from both Sides of the Equation, and it will be xx = aa + bb. Q. E. D.

AXIOM III. The Sum of the Hypothenuse and One of the Legs multiplied by their Difference, the Square Root of that Product will be the other Leg required.—This comes into use, when the Hypothenuse and One Leg is given, to find the other Leg.

This is only a Consequence of the last Demonstration; for it is there proved, that

aa + bb = xx. Consequently xx - bb = aa. But  $x + b \times x - b = xx - bb$ . Q. E. D.

AXIOM IV. Half the Longer of the Two Legs, added to the Hypothenuse, is always in Proportion to 86\*, as the Shorter Leg is to its opposite Angle.--This is useful, when the Sides are given, to find the Angles.

Construction.

Let CAE be the

Triangle proposed.

With the Radius CA

= the Hypothenuse,
describe the Semicircle IKDC, and,
producing the Diameter towards R, make DR = DC, the Radius. Draw the Tangent

IB and the Chord IA. Then will the Chord IA be equal to, or suf-

meter towards R, make DR = DC, the Radius. Draw the Tangent IB, and the Chord IA. Then will the Chord IA be equal to, or sufficiently near equal to IB, the Tangent of the Angle ACE; and RI is Triple of the Radius CI = the Hypothenuse CA.

#### DEMONSTRATION.

Supposing IB = IA; the two Triangles RIB, and REA are similar, because AE is parallel to BI. Therefore it will be (by the 6th Euclid's 4th.)

As RE : EA :: RI : IB = IA. And,

As RE : EA :: RI : IA = IB.

The Chord IA must be substituted for an Arch of a Circle whose Circumference = 360 equal Parts, whose Half is DKI. But the Radius of a Circle whose Circumference is 360, is 57.2 957 = CI, which tripled is 171.8871, or 172 nearly, = RI, by Construction; and RD = DC = CA, the Hypothenuse: Therefore it will be, as RE = RD + DC + CE, or twice the Hypothenuse, + CE the Base, is to RI, thrice the Radius = 172; so is EA, the shortest Leg to IB = IA = the Angle ACE, required. Or, more briefly, by taking Half the two sirst Terms, it will be,

As CA +  $\frac{CE}{2}$ : to  $\frac{RI}{2}$  = 86:: fo is AE the shortest Leg: to its opposite

Angle ACE required. Q. E. 1).

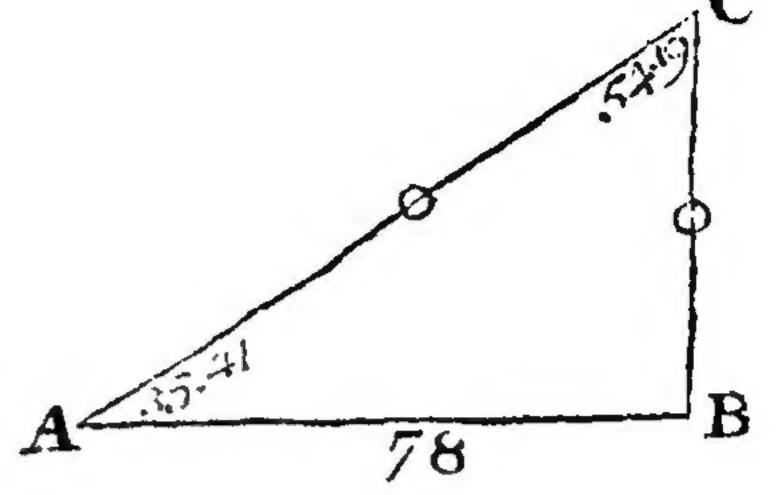
Note. These four Axioms will answer all the Cases of Right and Oblique Angled Triangles, except the last Case in Obliques, which will require some further Assistance, and will be shewn when we come to treat of that Case.

\* 86 = Radius and Half of a Circle whose Circumserence is 360, or, 85.94368.

C A S E

## CASE I.

The Acute Angles, and one Leg given; to find the Hypothenuse and the other Leg.

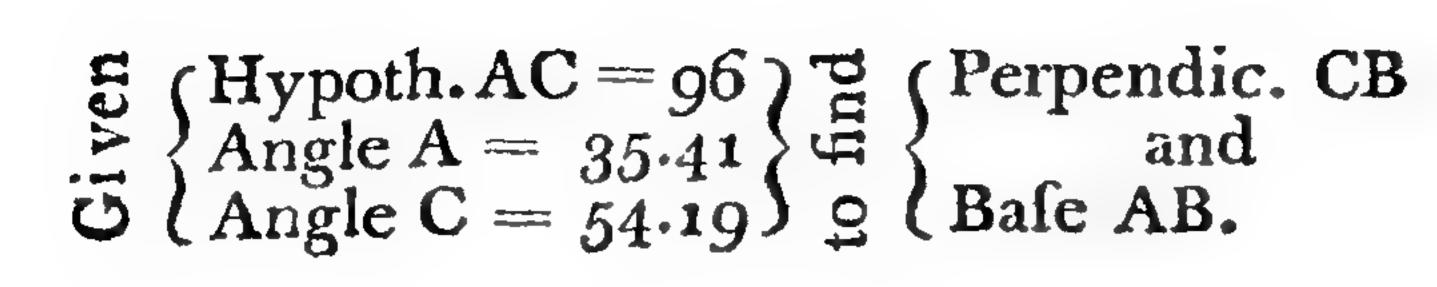


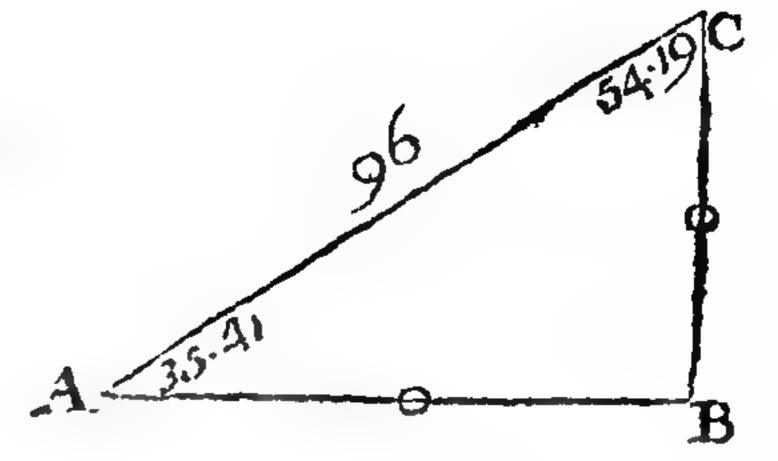
(1st.) Find the Natural Radius by Axiom I.

#### (3d.) Find the Perpendicular by Axiom III.

## C A S II.

The Hypothenuse and Angles given, to find the Two Legs.





#### (1st.) Find the Natural Radius by Axiom I.

## (2d.) Find the Perpendicular by Axiom I.

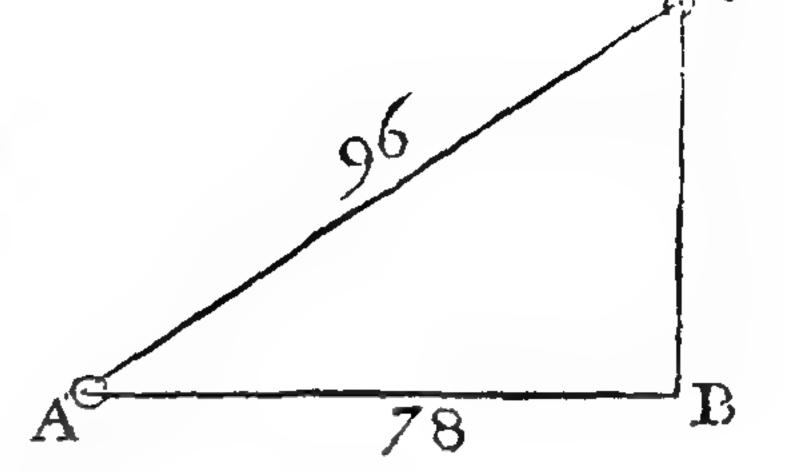
#### (3d.) Find the Base by Axiom III.

#### 12

## C A S III.

The Hypothenuse and One Leg given, to find the Angles and the other Leg.

Base AB = 78
Hypoth. AC = 96
Perpendic. CB
and
Angles A & C.



(1st.) Find the Perpendicular by Axiom III.

96 To Hypothenuse 78 Add the Base

174 Sum multiply 18 By Difference

1392

Extract the Root 3132(55.9 + Perpendicular

25 105)632 525 (109)10700 9981

(2d.) Find the Angle by Axiom IV.

To Hypothenuse 96 Add half longer Leg 39

Sum 135 : Fixed Number :: Perpendicular

56
86

336
448

135)4816.(35.67 + Angle A.
405

766
675

910
810

1000
945

Answer, {The Angle, 35" 41' nearly. The Perpendicular, 55.9+, or 56.

y 6 91,

## C A S E IV.

The Two Legs given; to find the Hypothenuse and the Angles.

(1st.) Find the Hypothenuse by Axiom II.

Answer, {The Hypothenuse, 96. The Angle, 35°.674. or 35° 41' nearly.

Note. Thus all the Cases of Right Angled Triangles, are easily and readily answered: and by the same Rules, and with the like Ease may the Oblique Angled Triangles be answered, as will evidently appear in the sollowing Cases.

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## Of Oblique TRIANGLES.

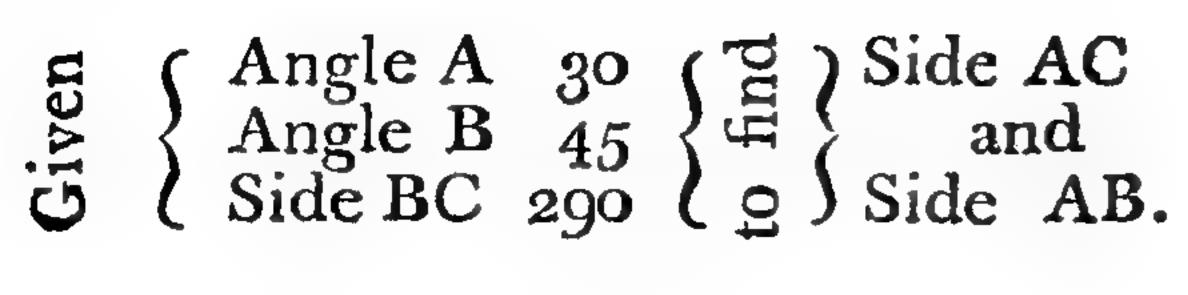
In the Solution of an Oblique Triangle, it is necessary, by this Method, to divide it into Two Right Angled Triangles, by means of a Perpendicular, which must always fall from the End of a given Side, and opposite to a given Angle.

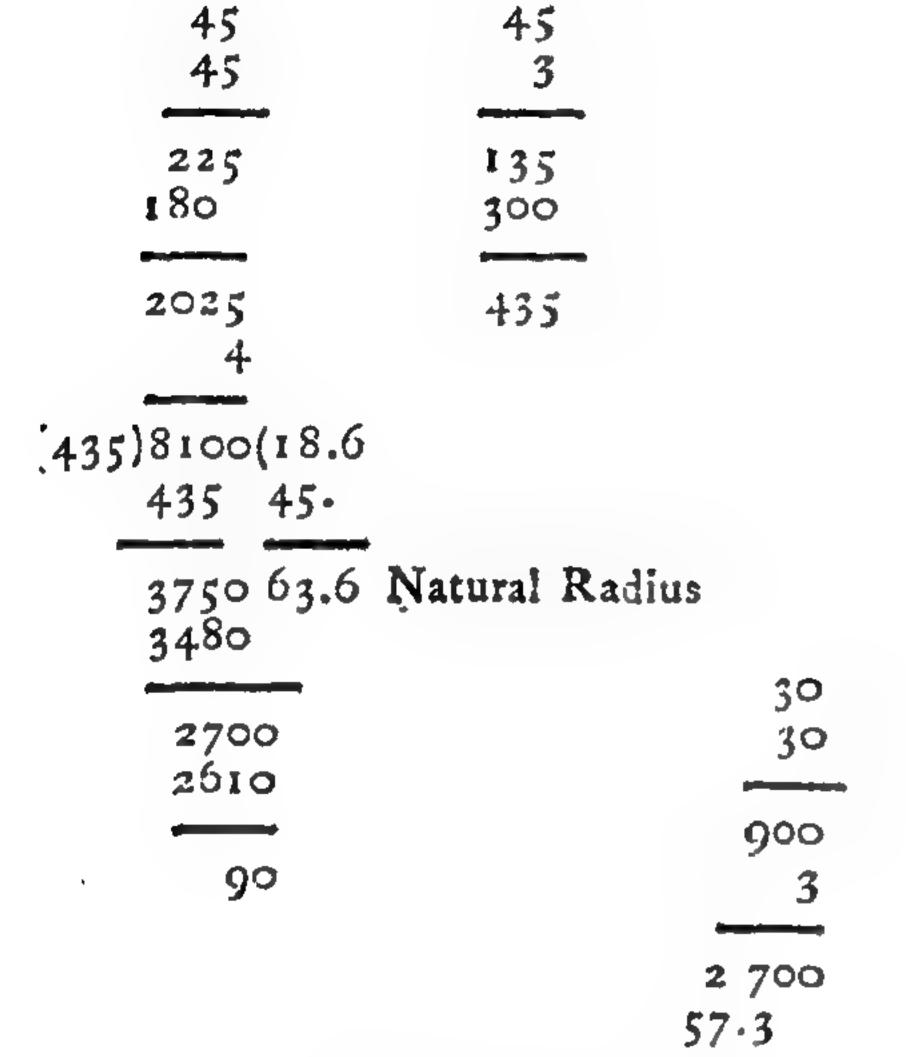
By this means the Perpendicular will sometimes fall within, and sometimes without the Triangle: When it falls within, it falls upon some Part of the Base, or longest Side; but when it falls without, it falls upon one of the shorter Sides continued. In either Case, there are Two Right Angled Triangles made, and the Angles, or Sides sought, are sound as if they were Parts of a Right Angled Triangle, by the soregoing Axioms; but it requires Two or Three Operations.

•

## CASE I.

Two Angles, and a Side opposite to one of them, given; to find the other two Sides.





Natural Radius 60.0

#### (3d.) Find the Side AD.

50

To Hypoth. AC 410
Add Perpend. CD 205

Sum 615

Multiply by Difference 205

12300

Extract the Root 126075(355 AD 9

65)360
325

705)3575
3525

B.

(1st.) Find the Perpendicular CD.

N. Rad. Hypoth. 
$$\angle$$
 at B

As 63.6 290 45

45

1450
1160

63.6) 13050(205 Perpendicular CD.
1272

3300
3180

# (2d.) Find the Side AC. As 30 Perp. CD N. Rad. As 30 60

30)12300(410 The Side AC Hypothenuse.
120
30
30

#### (4th.) Find the Side DB.

To Hypoth. CB 290
Add Perpend. CD 205

Sum 495

Multiply by Difference 85

2475
3960

Extract the Root 42075(205 DB.

4

405)2075
2025

Answer, { The Side AC 410. The Side AB 560. The Angle C 105.

#### 16

## C A S E II.

Two Sides, and an Angle opposite to One of them, being given; to find the rest.

Side AB 560 \ Side AC 410 \ Angle C and Angle B 45° \ Side BC. (3d.) Find the Angle DAC. To Hyp. CA 410 Add Ferp. AD 198 Fix'd Num. Side CD (1st.) Find the Side AD. As 608 -Hyp. BA 86 Nat. Rad. 工,B 63.6 As 560 45 636 45 848 2800 608)9116(14.9+, or 15° & A 2240 608 63.6,25200.0(396 AD 3036 1908 2432 6120 6040 5724 5472 3960 568 3816 (4th.) Find Side BD. 144 To Hypoth. BA 560 Add Perpend. DA (2d.) Find CD. Sum 956 To Hypoth. CA 410 Multiply by Differ. 164 Add Perpend. AD 396 3824 Sum 806 5736 956 Multiply by Difference 3224 Extract the Root 156784(395.9 + or 396 BD 805 Extract the Root 11284(106 CD 69 667 206(1284 785(4684 1236 3925 7909)75900

Then from BD = 396 Take --- CD = 106

Hence we find by Inspection.

The Angle ACD = 75

The Angle ACB = 105

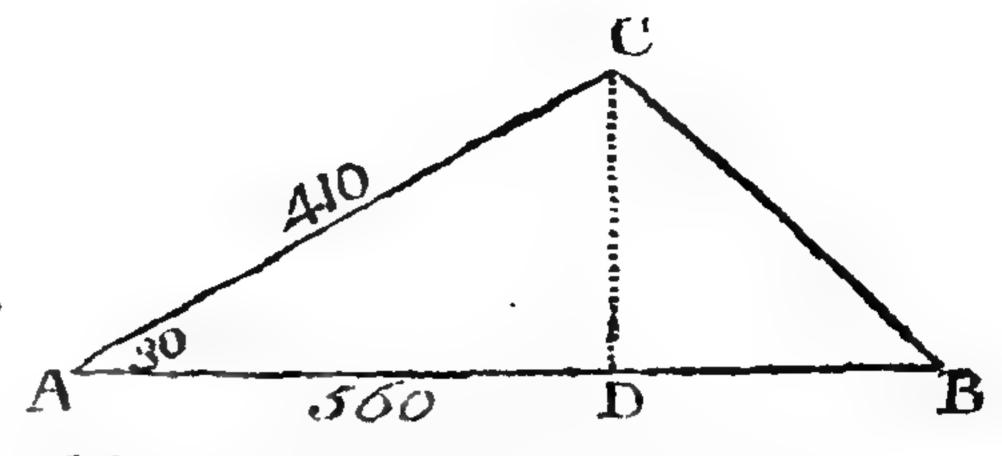
Remains — BC = 290 Required. and The Angle BAC = 30° Requ.

In this, and several of the following Cases and Problems, the Operation for finding the Natural Radius is omitted, for Want of Room.

## C A S III.

Two Sides, with the Angle comprehended by them, given; to find the rest.

Side AC 410 ) Side BC and Side AB 550 Side AB 560 Angles B and C.



(1st.) Find the Perpendicular CD.

(3d.) Find the Side BC.

BD squar'd 205
205 CD squar'd 205
205
1025
4100
4100

To Square of DB 42025
Add Square of CD 42025

Extract the Root 84050(280.0 +. or 200 BC

Extract the Root 84050(289.9 +, or 290 BC

4

48)440
384

569)5650
5121

5789)52900
52101

(2d.) Find the Part AD.

## (4th.) Find the Angle B.

799

From AB = 560 Take AD = 355 Remains BD = 205

The Angle A = 30, added to Angle B = 45, and then subtracted from 180, leaves 105 for Angle C.

Answer, { The Side BC 290 The Angle B 45° The Angle C 105°.

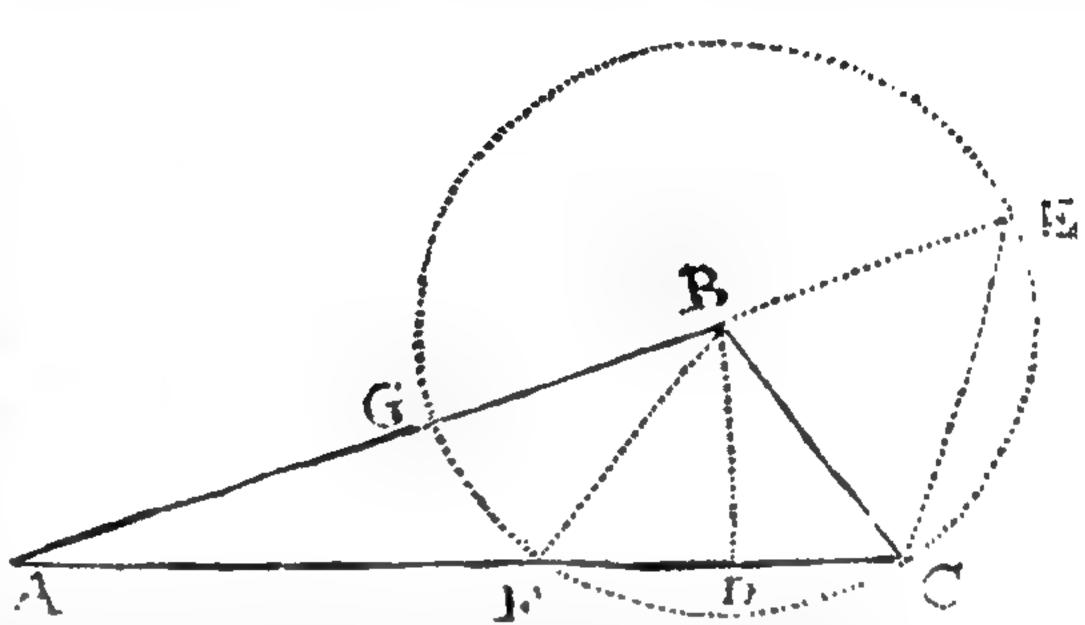
CASE

The Three Sides given; to find the three Angles.

In all Triangles, as the Base or greater Side is to the Sum of the other Two Sides; so is the Difference of the Sides to the Difference of the Segments of the Base; which Half Difference, added to Half the Base, the Sum will be the Greater Segment, upon which the Perpendicular falls: But if subtracted from Half the Base, the Remainder will be the Less Segment: So will the Oblique Triangle be reduced to two Right Angled ones, and may be answered after the same Manner as before.

CONSTRUCTION.

Let ABC be the Triangle proposed, whose three Sides AB, AC, and BC are given. On B, as a Centre, with the Radius BC, describe the Circle GFCEB; continue the Line AB till it cuts the Circle in E; let fall the Perpendicular BD; draw the Lines EC and BF. Then will AE

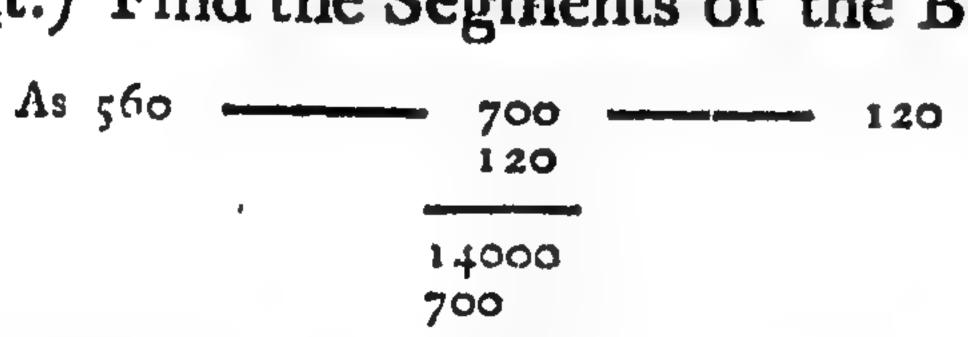


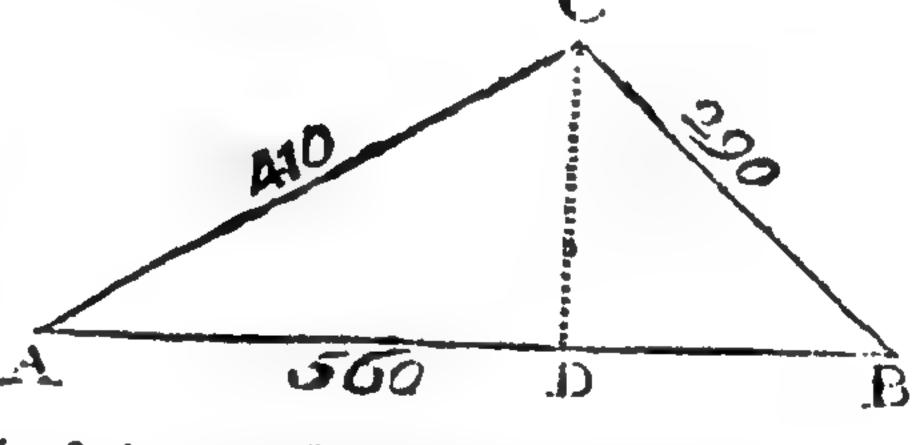
be the Sum of the two Sides AB and BC, and AG, their Difference. Also AC = AD + DC, the Sum of the Segments of the Base, and AF = AD — FD, their Difference.

Demonstration. Because the Triangles ABF, and AEC, are similar, it will be, as AC: AE:: AG: AF. Q. E. D.

Side AB 560 ( ) Angle A Side AC 410 ( ) Angle B Side BC 290 ( ) Angle C.

(1st.) Find the Segments of the Base.





(3d.) Find the Angle at A.

To 410 Side AC Add 177 half AD 560)84000(150 Disserence Fix'd Num. Side DC To ! Base 280 As 587 — 86 — 205 — 30 \( \text{A} 560 Add E Diff. 75 2800

G.Scg.AD 355 2800 From Base 280 Subtract & Diff. 75

Lesser Seg. DB 205

(2d.) Find the Perpendicular CD, by Case III. = 205.

(4th.) Find the Angle at B.

To 290 Side BC Add 102 half DB - Fix'd Num. Side DC As 392 --- 86 -- 205 -- 49.9+, or 45 4 B

Then the Angle A = 30°; added to B = 45°, and subtracted from 180°, leaves 105° for the Angle C, which were the Angles required.

THE

#### THE USE OF

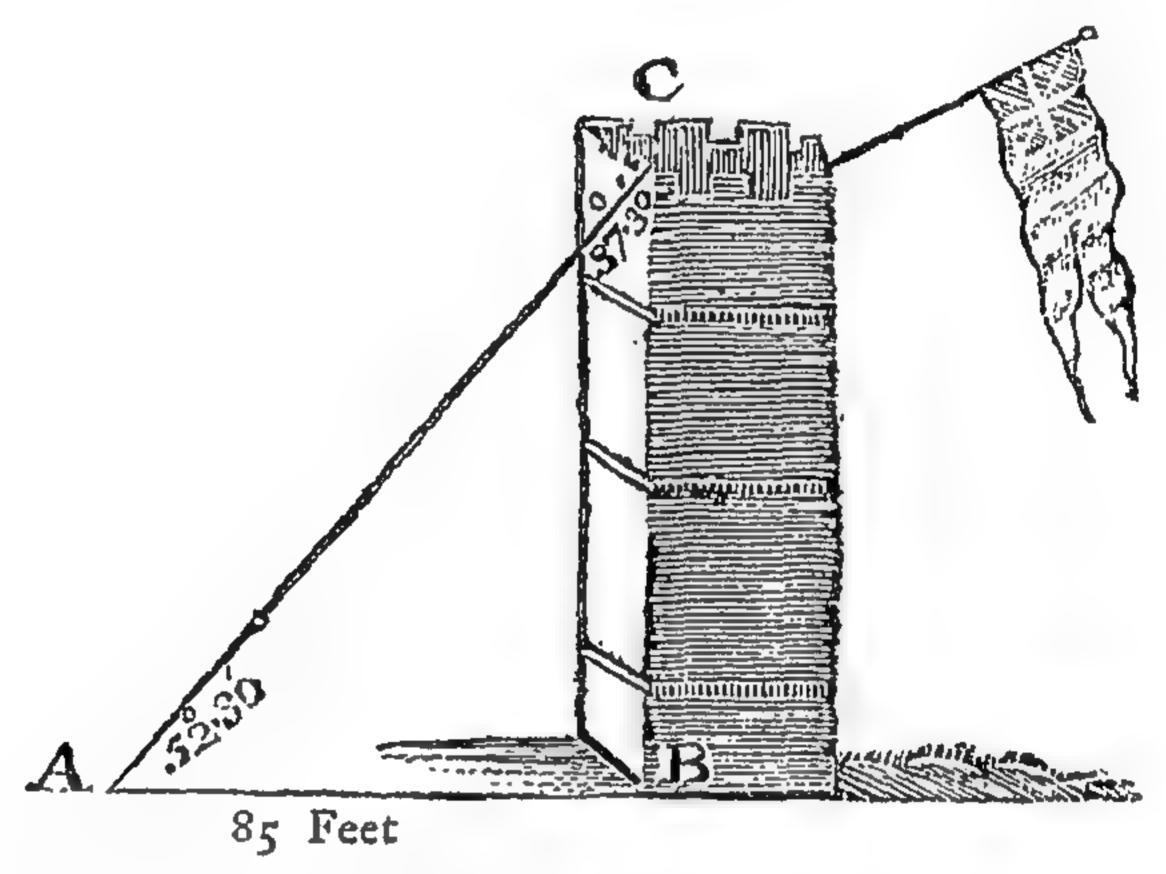
## TRIGONOMETRY

Exhibited in the Solutions of a Number of interesting Problems; many of which every Day occur; are of the greatest Utility in the Army, Navy, &c. and cannot be answer'd without it.

#### PROBLEM

To take the Height of any accessible Object at one Station.

First, with a Quadrant, by looking through the Sights to the Top of the Tower, find the Quantity of the Angle A, which suppose 52° 30'; then measure the Distance AB, which suppose to be 85 Feet; from hence you may proceed to find the Height, by Case I. of Right Angled Triangles.



### (2d.) Find the Perpendicular BC.

To Hypothenuse 139.44 Add Base 85. (1st.) Find the Hypothenuse AC. Sum 224.44 Angle C: Base:: Nat. Rad. Multiply by Difference 54.44 As 37.5 - 85 - 61.52 89776 89776 30760 89776 49216 112220 Extract the Root 12218.5136(110.537 Perpend. 37.5)5229.20.(139.44+Hyp. 375 1479 21)22 1125 .2205)11851 3542 3375 11025 1670 22103)82636 66309 1500 221067)1632700 1700 1500 1547469

Answer, 110.537 +, or 110 Feet, and above \(\frac{1}{2}\): The Height required.

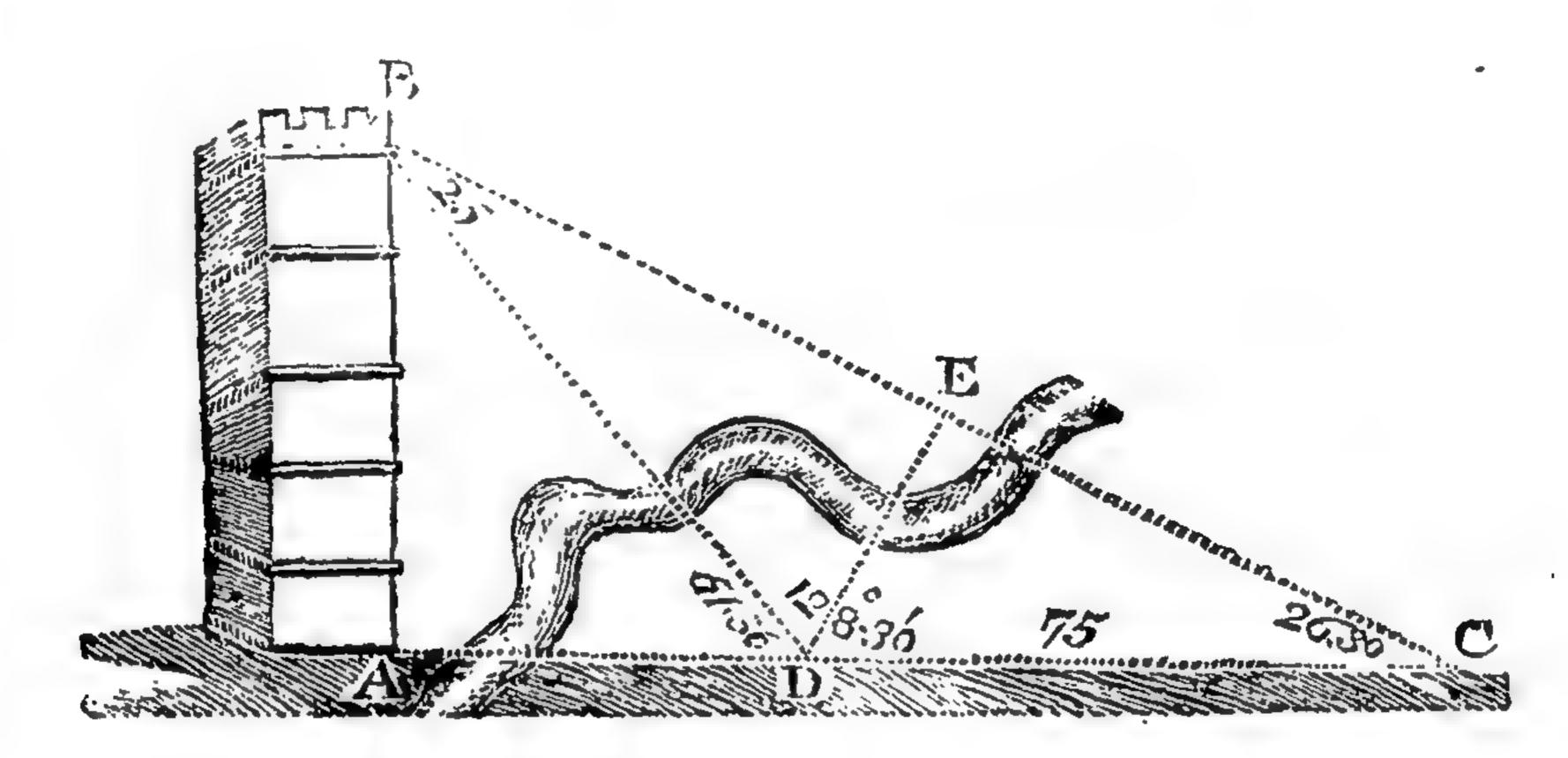
Note: That in this, and all such Cases, you must add the Height of your Eye, or Instrument, to the Altitude before found.

## PROBLEM.

To measure an Inaccessible Altitude.

Let AB, in the following Figure, be a Church, Tower, or Fort, whose Height is required; but by Reason of a River, or some other Obstacle, it is inaccessible; that is, you cannot come to the Foot of it, by Reason of the Water, &c.

First, with a Quadrant, take the Angle of Altitude at C, which suppose 26° 30°. Then incasure in a Right Line towards the Tower to D, any Distance, suppose 75 Feet, and at D observe again the Angle of Altitude, which let be 51° 30′.



Then; the two Visual Lines CB and DB, with the Distance DC, make the Oblique Triangle CBD, in which are given—All the Angles and Side CD. The Angles BCD being 26° 30′, and the Complement of ADB 51° 30′ to 180, is the Obtuse Angle BDC 128° 30. Consequently, the third Angle CBD, at the Top, is = 25°.

(1st.) Find the Perpendicular DE in Triangle DBC.

Nat. Rad.: Op. Side DC:: Ang. C: Perp. DE As 59 4 - 75 - 20.5 - 33.46

(3d.) Find the Height AB in the Right Angled Triangle ABD.

Nat. Rad.: Op. Side BD:: Ang D: Height As 65.7 — 79.19 — 51.5 — 62 AB

(cd.) Find the Visual Line BD in Triangle BDE.

Ang. B: Op. Side DE:: N. Rad.: Side BD As 25 — 3340 — 59.17 — 79.19 (4th.) Find the Distance AD in the Triangle ABD.

N. Rad.: Op. Side BD:: Arg ABD: Dif. AD
As 61.7 79 19 - 38.5 49 41

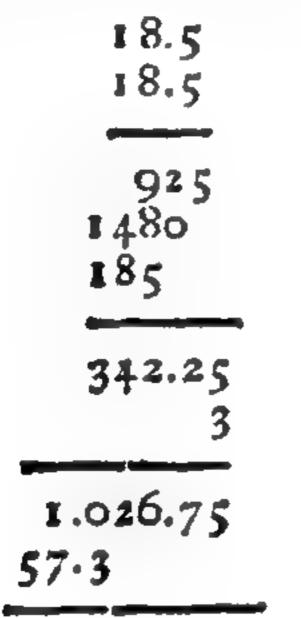
Nors: The Line BD is the Length of a Scaling Ladder, which would reach from the Station at D over the Fojs of Elich, to the Top of the Tower at B.

## PROBLEMII.

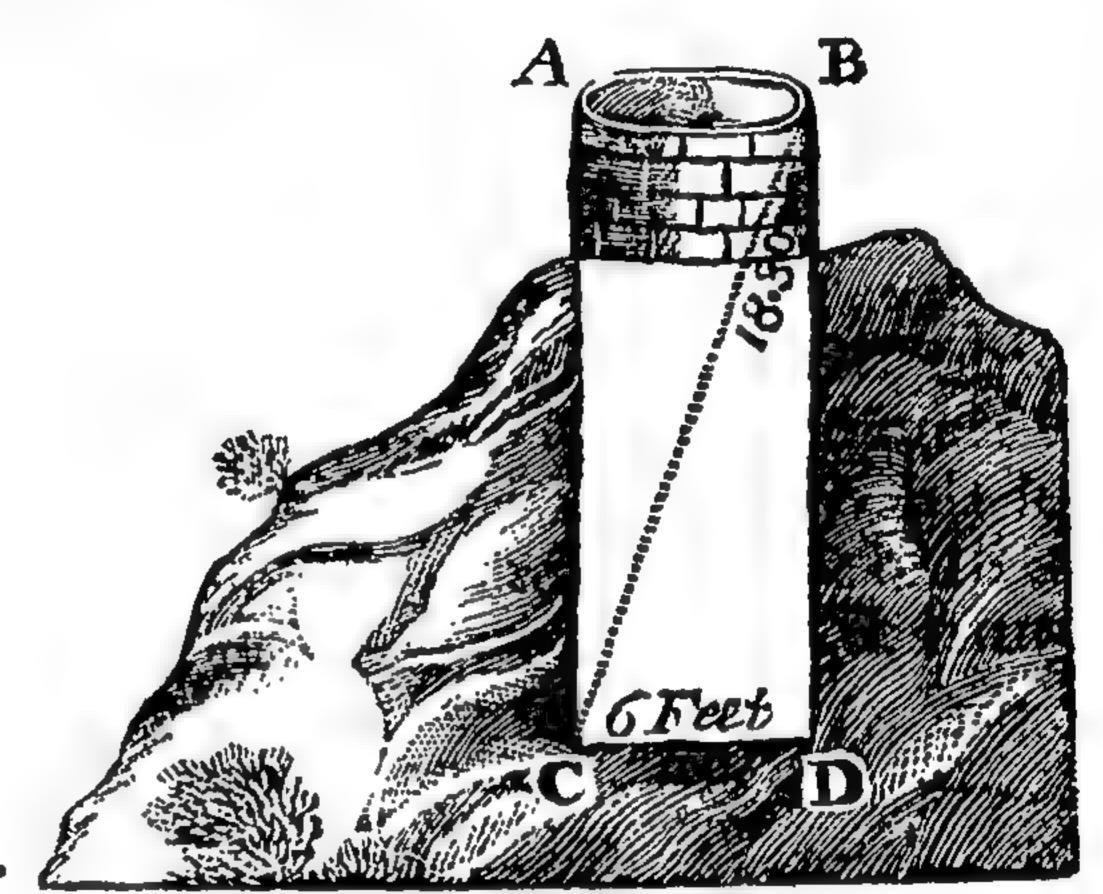
To measure the Depth of a Well, or the Height of an Object from the Top of it.

First, look through the Sights of the Quadrant to the Bottom of the opposite Side the Well at C, so you will have the Angle CBD; next, take the Breadth AB at the Top, which is equal to CD at the Bottom: Then, by Case I. of Right Angle Triangles, you may easily find the Depth BD required.

Suppose the Angle at B, by Observation, to be 18° 30', and the Breadth at the Top 6 Feet: What's the Depth?



Natural Radius 58.32675 but 58.3 is enough.



(1st.) Find the Hypothenuse BC.

As 18.5 Op. Side N. Rad.

18.5)349.8.(18.9 + the Line BC)

185

1648

1480

1680

1665

(2d.) Find the Depth BD.
To Hypoth. BC 18.9

Add Side CD 6.

C Sum 24 9

Multiply by Differ. 12.9

2241
498
249

Extract the Root 321.21(17.89 the Depth AB

27)221
189

Answer, { 17.89 +, or 18 Feet, the Depth required.

15

348)3221 2784 3569)43700 32121

## PROBLEMIV.

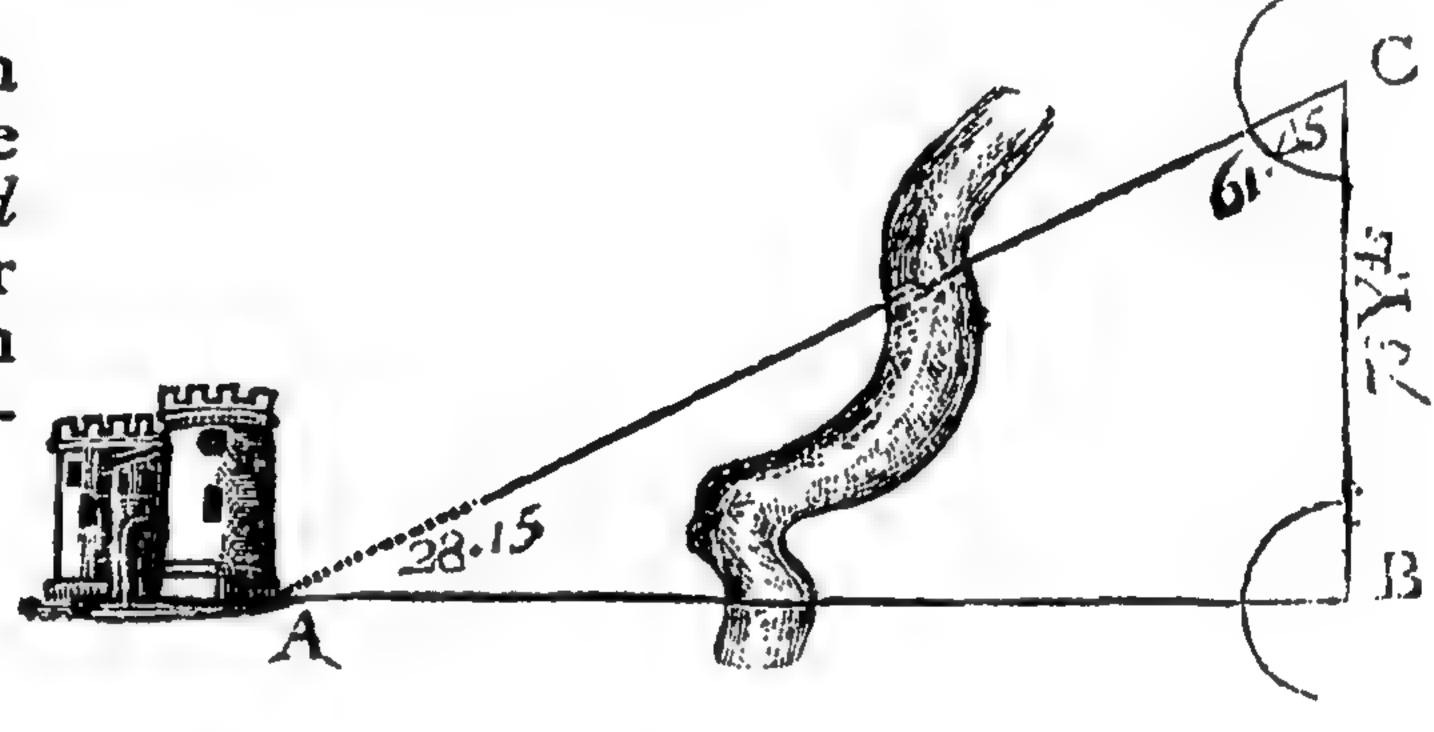
To measure the Distance of any Object.

Suppose yourself standing at B, and a great Way off, as at A, you see a Fort or Castle, &c. or any other Object, whose Distance you would

find from the Place where you now stand.

First, a Theodolite, or Semicircle, being placed at B, lay the Index, with its Sights, on the Diameter, where the Degrees begin, and through them view the Castle, &c. at A. The Instrument remaining fix'd in this Position, move the Index to 90 Degrees, (that being a Right Angle) and view some Mark at a Distance, (the farther off the better) as at C.—Next measuring the Distance from B to C, which suppose 73 Yards, remove your Instrument, and set it up at C. Then, with the Index laid upon the Beginning of the Degrees, as before, turn the Instrument about, till you can see your first Station at B, where fasten it; then turn the Index till you can see the Object A, and observe what Degrees are cut, as suppose 61° 45', which is the Quantity of the Angle where you stand; whose Complement to 90° is the Angle A.

Now, here are given all the Angles, and one Side of a Right Angled Triangle, to find either of the other Sides, which will be the Distance required.



(2d.) Find the Distance from B.

To Hypoth. AC 154.2
Add Side BC 73

Sum 227.2

Multiply by Differ. 81.2

4544
2272
18176

(1st.) Find the Distance from C.

Extract the Root 18448.64(135.8 +, Dist. from B

## PROBLEM.

To take the Distances of several inaccessible Objects, as Forts, Churches, in a Town, or Squadron of Ships at Sea, and to delineate them upon Paper.

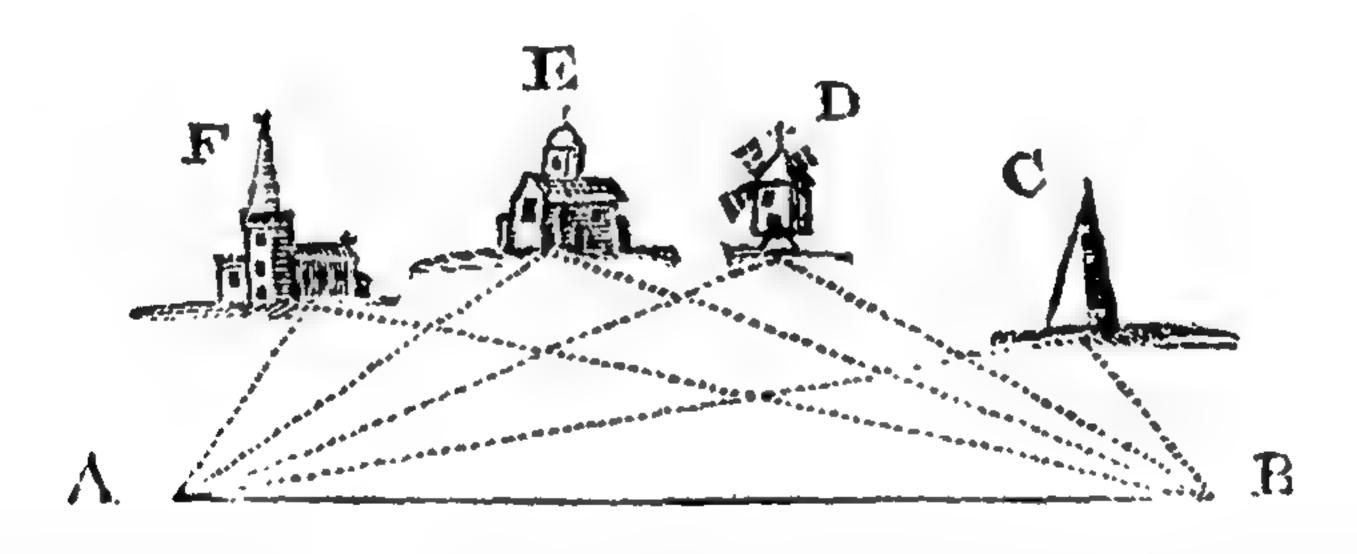
First, make choice of two places, from either of which you may conveniently see all the Objects; which two Places let be A and B in the following Figure.—This being done, set up your Instrument at A, laying the Index on the Diameter, and turn the whole Instrument about, till, through the Sights, you see your second Station B. Then, fixing the Instrument, direct your Sights to the several Objects, C, D, E, and F; noting down the Degrees cut at each Observation, which suppose to be as in the Table.

Then, remove the Instrument to B, laying the *Index* on the *Diameter*, and turn it about, till, through the Sights, you see your former Station at A; then direct your Sights to every one of the Objects at C, D, &c. setting down

the Degrees at each Observation, as in the Table. Also measure the Stationary Distance, and set that down.

	C	D	E	F
tst Stati on	1 1	23	36	59
2d Station	51	31	22	13
Stationary E	diftan	ce 15	o Yar	ds.

First, upon a Piece of Paper draw the Line AB; and from a Scale of equal Parts, take off, with your Dividers, the Stationary Distance = 150, and set it from A to B, so will A represent your first Station, and B the second. Then lay the Center of the Protractor upon the



Point A, with its Diameter upon the Line AB; keeping it fast, make Marks by the Edge at 11, 23, 36, 59, and draw Lines from the Point A through each of those Marks. Then upon B place the Center of your Protractor, its Diameter lying upon the Line AB; make Marks by the Side at 13, 22, 31, 51. Then draw Lines from the Point B through each of these Marks, and where the Lines cut the former correspondent Lines there will be found the Places representing these Objects. Then any of these Lines being taken in a Pair of Dividers, and applied to the Scale you laid your Stationary Distance down by, will give you their Distances, either from your Stations or from one another.

The Distance of any of these Objects from either Station, &c. may be found by Calculation; one Side and the Angles being given: But I shall omit that, on purpose to exercise the Learner's Genius, and proceed.

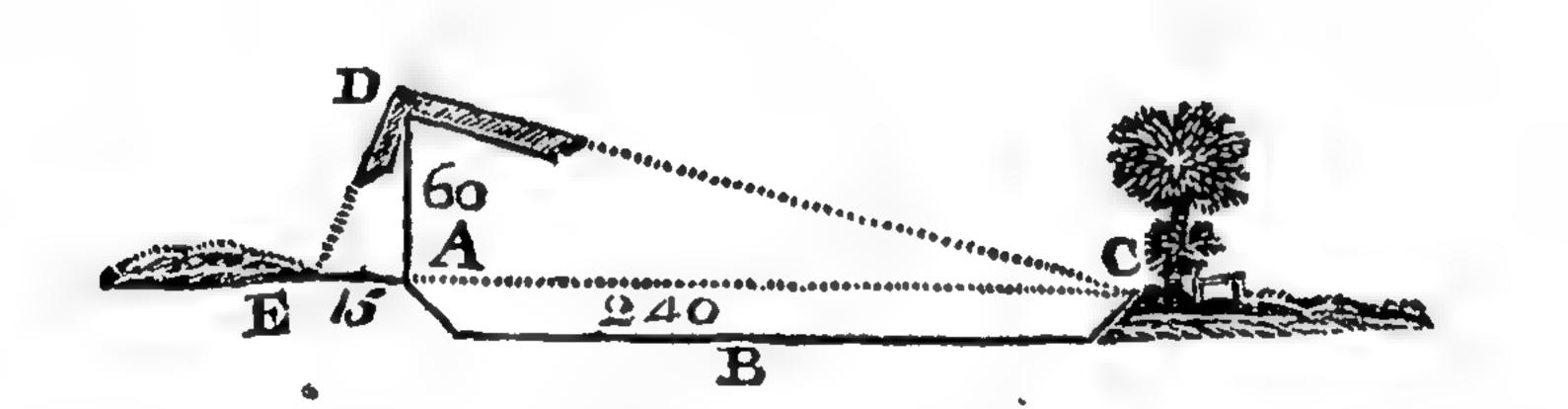
## PROBLEM.

To take the Distance, upon level Ground, of any inaccessible Tree, Fort, &c. or Breadth of a River, by a Common Square.

Suppose there is a River, as ABC, whose Breadth you want to know.—First, upon the Bank, at A, set up a Stick, AD, which suppose to be 5 Feet, or 60 Inches high; then fixing your Square on the Top, at D, look by the Side of it till you see the Edge of the opposite Shore C, and fasten it, as it may not go from that Position. This done, extend a Thread from D, by the other Side of the Square till it touch the Ground at E. Then measure the Distance EA, which suppose 15 Inches, (or 1 Foot 3 Inches) and you may find AC (by Reason of similar Triangles) thus.

Dist. EA: Side DA: Side DA: Dist. AC

As 15—60—60—240 Inches,
which, reduc'd to Feet, give 20 for the Breadth of the River sought.



Note. There are various Ways of taking Heights and Distances; but the best is to take the Angles for Heights by a Quadrant; and the Angles for Distances by a Semicircle or Theodolite; and calculate by the foregoing Axioms. In all Heights the Triangle stands upright; but in Distances, it is supposed to lie stat or horizontal.

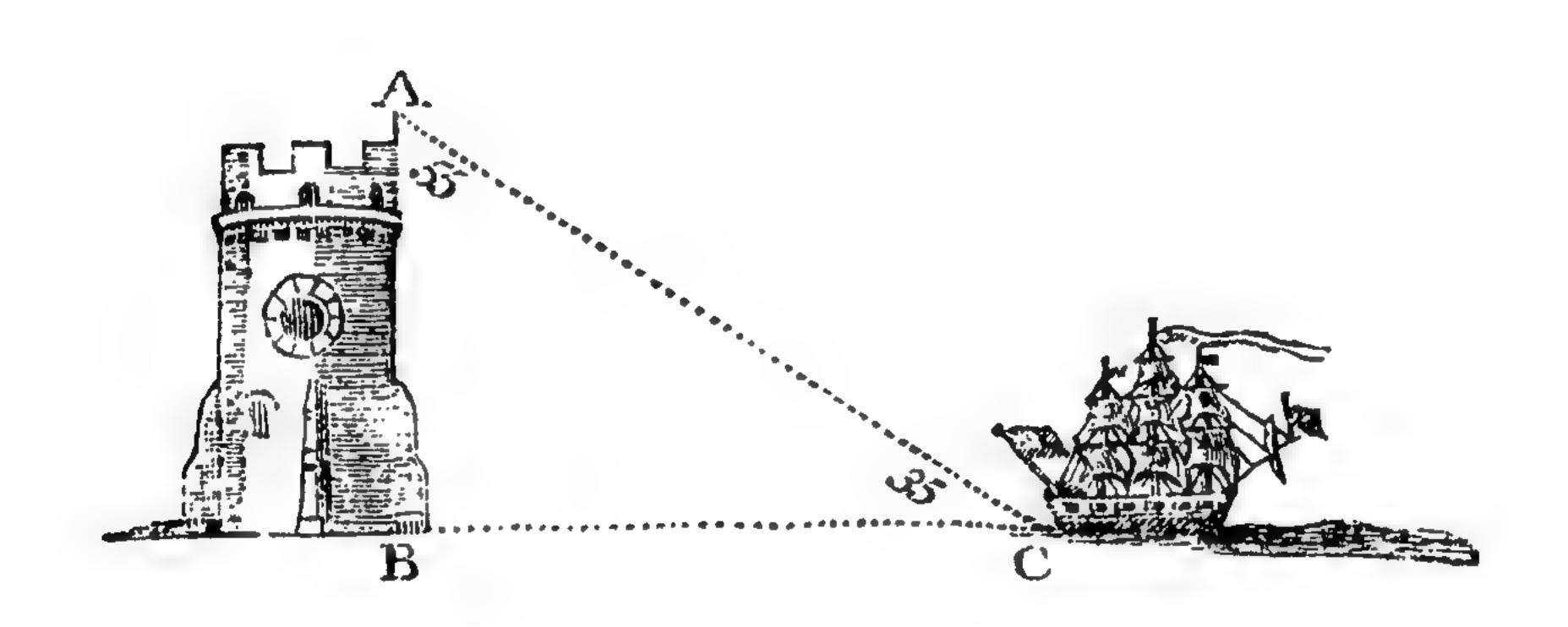
River, or a small Distance, without any Instrument whatever; which is thus. Standing upon the Bank, bring down the Edge of your Hat, till it appears to touch the opposite Side, then steady your Head by laying your Hand under your Chin, and turn yourself towards some level Ground, observing where the Edge of your Hat glances upon it; for then, the Distance from you to that Place, is equal to the Breadth of the River, or Distance required.

## PROBLEMUII.

To find, from the Top of a Fort, or Tower, how far any Tree, Ship, &c. is from you.

Let A be the Top of a Tower or Castle standing by the Sea Side; and let C be a Ship at Sca, or lying at Anchor, and you would know how far that Ship is off the Castle Wall.

With your Quadrant or Semicircle, direct your Sights from the Top of the Tower to the Place where the Ship is, and take the Angle, which we will suppose to be 55 Degrees. Then the Castle Wall being known before to be 143 Feet high, you may easily find the Distance of the Ship from the Wall in this Manner.



(1st.) Find the Side AC.

(2d.) Find the Base BC.

By this Method you may easily discover if a Fleet of Ships, or one fingle Ship, at Sea, makes towards you or not. For having observed from the Top of the Fort the Angle from thence to the Ship, and noted it down, rest a little Time, and observe again: Then, if the Angle be bigger than before, the Ship is departing from you; but if less she is making towards you.

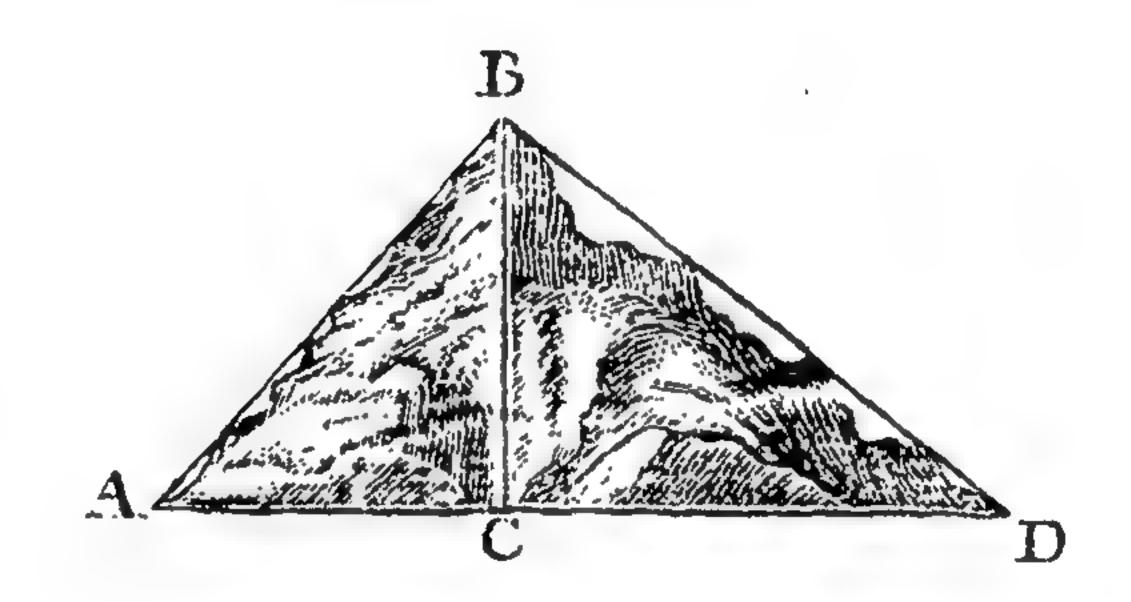
## PROBLEM VIII.

To take the Perpendicular Height of a Hill or Mountain, and also the Horizontal Line, or Base, on which it stands.

Let ABD be the Hill.—First, set up a Mark on the Top at B, equal to the Height of the Quadrant or Instrument that is used at the Bottom, from whence you intend to make your Observation. Then by looking through the Sights to B take the Quantity of the Angle at A, which we will suppose to be 50°. Next measure the Hill from A to B, which let be 546 Feet. This being done, you may easily find the Perpendicular BC, or Part of the Base AC, by Case II. of Right Angles.

#### For the Perpendicular BC.

For the Side AC.



Now, as the Hill descends, you may go on the opposite Side, and make the like Observations, viz. set up the Instrument at D, and take the Angle D, 40°, and measure the Side DB, 651 Feet, then you may find the Side CD in the same manner you did AC. Thus,

If to the Part AC = 351.1 Feet, be added the Part CD = 499 Feet, the Sum 850.1 Feet will be the whole Length of the Horizontal Line AD requir'd.

The Perpendicular Height DC is = 418.7 as above.

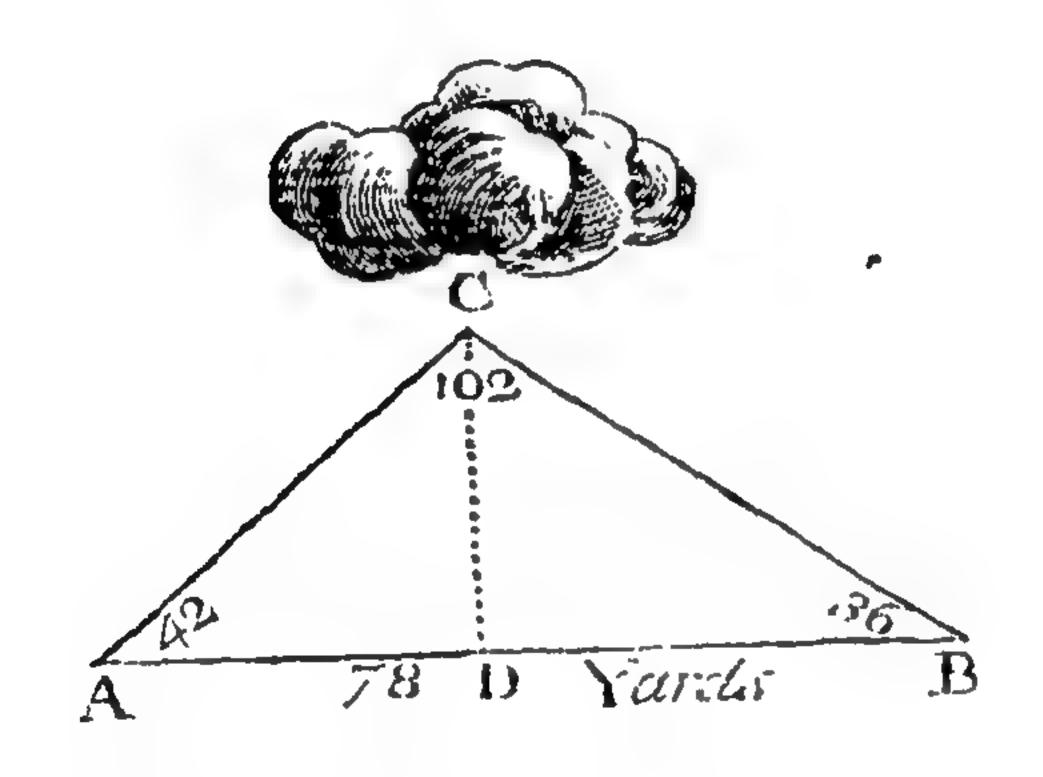
## PROBLEMIX.

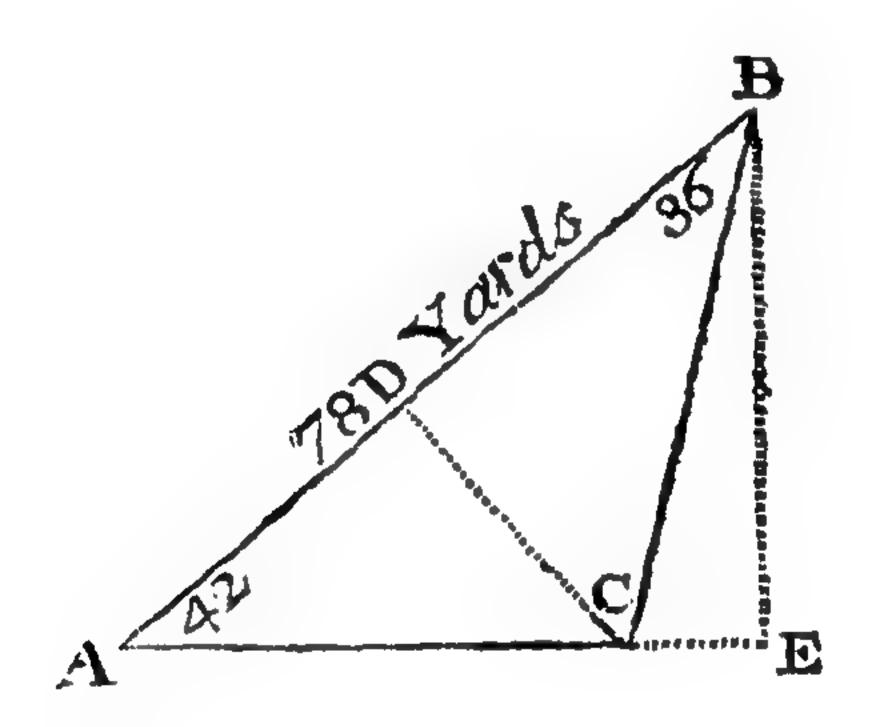
To take the Height and Distance of a Cloud.

Suppose it was required to find the Height of the Cloud C.

Let a Person standing at A, look through the Quadrant to the Cloud at C, so will the Thread cut the Angle at A. At the same Time let another Person, making the like Observation at B, take the Angle B. Then measure the Distance between the two Stations. By this means you will have one Side and all the Angles of an Obsique Angle Triangle given, from whence you may easily obtain the rest, and particularly the Perpendicular CD, which will be the Height of the Cloud required.

EXAMPLE. Suppose the Angle at A, by Observation, be 42°, the Angle B 36°, and the Distance AB 78 Yards: I demand the Height of the Cloud.





(1st.) Find the Perpendicular BE in Triangle ABE \*.

(2d.) Find the Hypothenuse CB in Triangle CBE.

N. Rad.: Op. Side AB.:: Ang. A: Perp. BE
As 62.6 — 78 — 42 — 52.3

Ang BCE: Op. Side BE:: N. Rad.: Hyp. BC
As 78 --- 52.3 --- 79 --- 52.96

(3d.) Find the Perpendicular CD in Triangle CDB.

N. Rad.: Op. Side BC:: Ang. DBA: Perp. CD
As 61.1 \_\_\_\_\_ 52.96 \_\_\_\_\_ 36 \_\_\_\_ 31.2

Answer, 31.2 Yards, the Height required.

The Figure on the Right Hand is only that on the Lest set in a different Position, to shew in a more natural or easy Manner, how the Perpendicular salls from the End of the given Side AB, upon the Side AC produced to E.

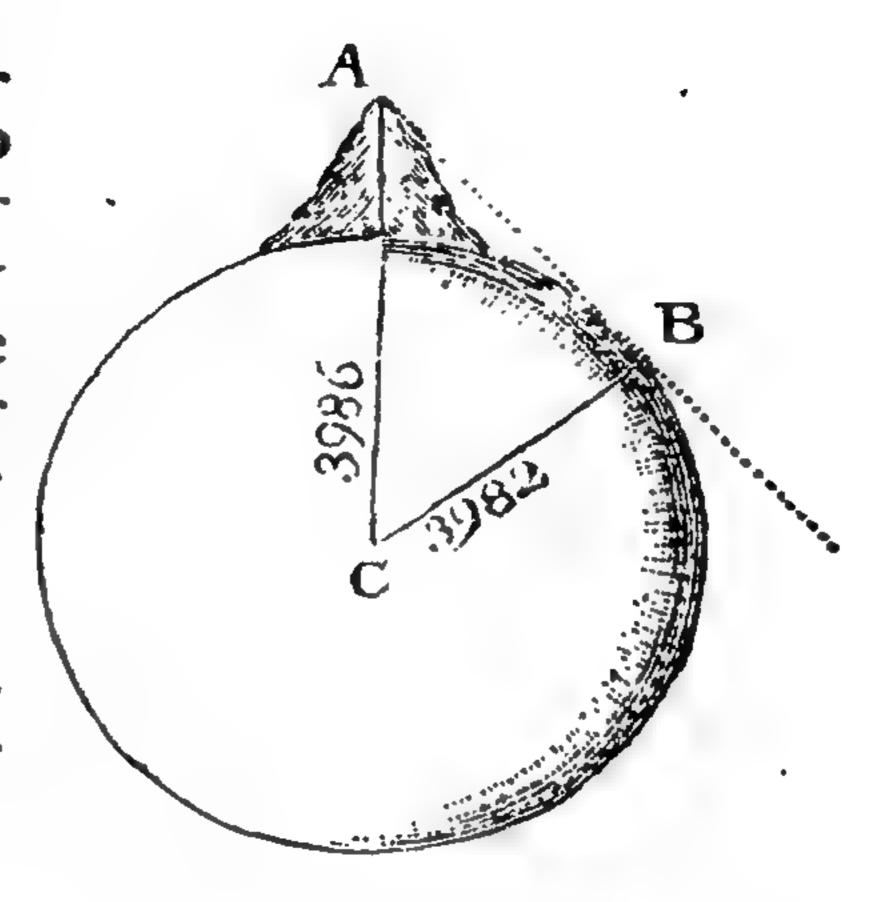
PROBLEM

## PROBLEM.

To find how far a Hill of any given Height can be seen at Sea, or upon level Ground.

How far, for Instance, can the Pike of Teneriff be seen at Sea, whose Height is about four Miles.

The Circumserence of the Earth is suppos'd by Mathematicians to be divided into 360 equal Parts, called Degrees; and our countryman, Mr. Norwood, has found, by measuring from the Tower of London to the Middle of the City of York, in the Year 1635, that one of those Degrees, upon the Earth's Surface, contains 692 Miles; according to which Measure, we find the Earth's Circumference to be 25020 Miles—its Diameter 7964—and its Semidiameter 3982.



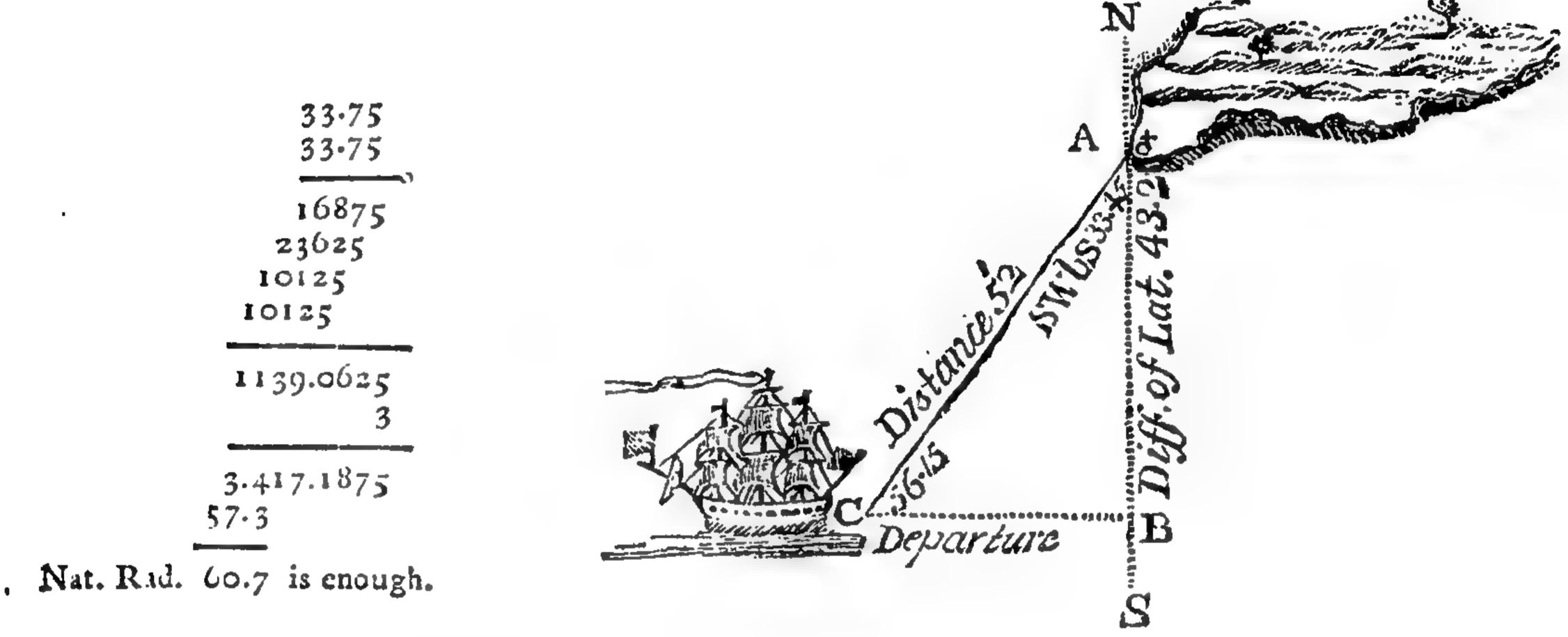
Then in the Triangle ABC, Right Angled at B, we have the Side CB = the Earth's Semidiameter 3982. Also the Line AC = the Semidiameter and Height of the Mountain together = 3986. To find AB, the Distance from the Hill to the visible Horizon.

This Mountain can be seen 178.5 Miles at Sea.

## PROBLEM XI.

The Distance run at Sea, and the Course, given; to find the Difference of Latitude and Departure from the Meridian.

Suppose a Ship from A, in the Latitude of 50° North, sails away, SW by S. 52 Miles, to C: I demand the Latitude she is in, and also her Departure from the Meridian.



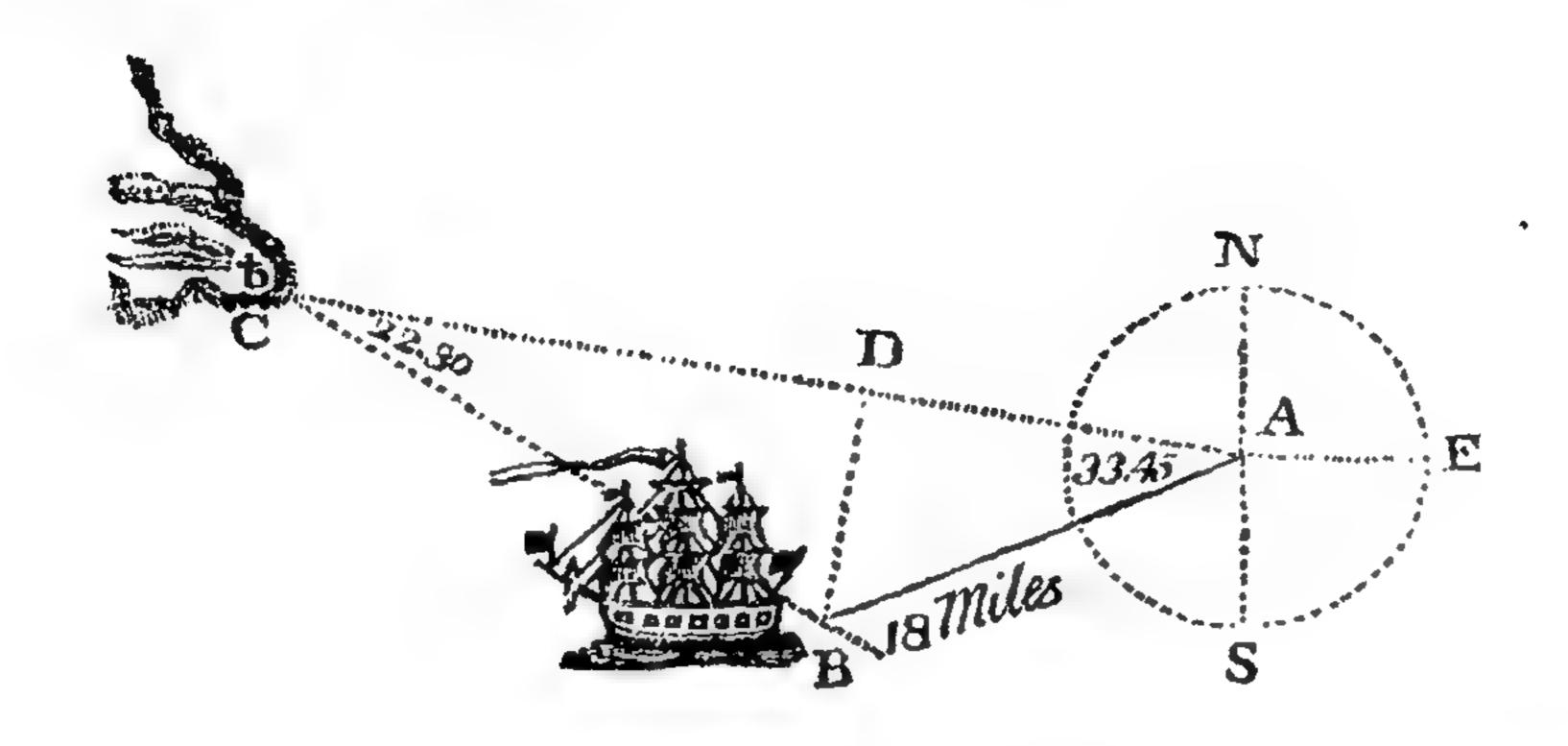
(1st.) For the Departure.

Note. That in all Cases of Sailing, we suppose the Top of the Book North, and Bottom South; the Right Hand East, Left Hand West. The Distance run is the Hypothenuse; the Dissertine of Latitude is the Perpendicular; the Departure the Base. The Angle at the Perpendicular is the Course, and the other its Complement.

## PROBLEM XII.

To take the Distance of any Cape, Fort, or Island, from a Ship at Sea.

Sailing W. S. W. I saw, at some Distance, a Point of Land, which I fet, and find it bears from me W. by N. and having sailed 6 Leagues further, I find it then bears from me N. W. by W. I would know how far this Land is from me.



33.75 33.75 16875 23625 10125 10125 3 3.417.1875 57.3

Nat. Rad. 60.7 is enough.

22.5 22.5 1125 450 450 506.25 3 1.518.75 57.3 N. Rad. 58.8 is enough? (1st.) Find the Perpend. BD in Triangle ABD.

N. Rad. Op. Side AB 
$$\angle$$
 A
As 60.7 — 18 — 33.75
18

27000
2375
60.7)607.50(10 Perpend.
607

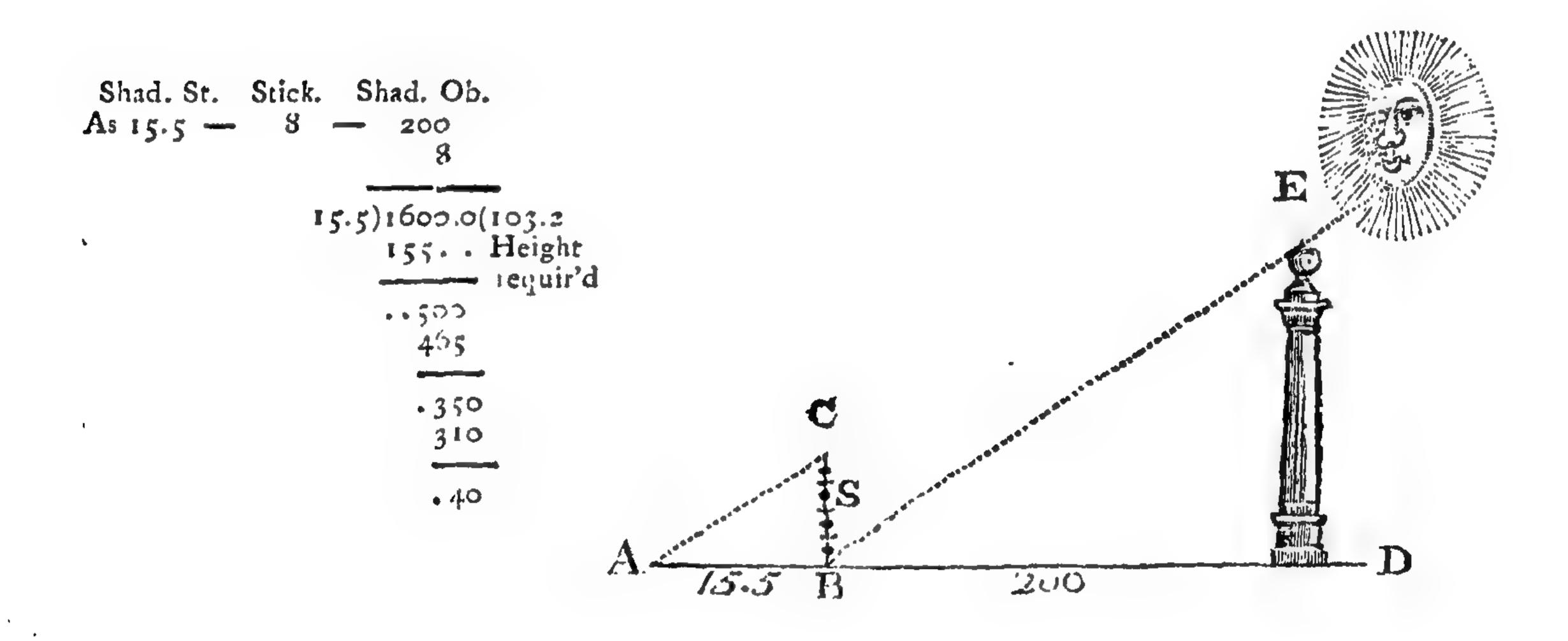
(2d.) Find the Distance CB in Triangle BCD.

Answer, 26.13 Miles, the Distance required.

## PROBLEM XIII.

To take the Height of a Tree, Fort, Obelisk, Pyramid, or any Object, by a common Stick only, when the Sun or Moon shines upon it.

Take a Stick of any Length, suppose 8 Feet; set it upright upon the Ground, as at CB in the Figure below. Mark the End of its Shadow at A, and measure its Length from B to A, which suppose to be 15.5 Feet. Then measure the Length of the Shadow of the Pillar or Obelisk BD, which suppose to be 200 Feet. This being done, you may easily find the Height: For (by Reason of like Triangles) it will always hold,—as the Length of the Shadow of the Stick AB in the small Triangle, is to its Height CB; so is the Length of the Shadow of the Obelisk BD in the great Triangle, to DE the Height thereof.



In this Manner the Heights of the Pyramids in Egypt have been taken. Those stupendous Buildings are supposed to have been erected by the Children of Israel, when in Bondage, for Sepulchers for the Egyptian Kings. They are the greatest Pieces of Antiquity now in Existence. There are several smaller Ones, but the largest, which is justly esteemed one of the Wonders of the World, is 500 Feet in Perpendicular Height;—700 Feet if measured obliquely from the Bottom to the Top;—and its Base covers about 11 Acres of Ground.

the form of the first the first

## PROBLEM XIV.

To take the Height of any accessible Object by a Bason of Water, or common Looking Glass.

Travelling along the Road, I see a fine May Pole, whose Height I would gladly know; but having no Mathematical Instrument with me, I procure a Bason of Water, which I set upon the Ground, at some Distance from the Pole, as at A; then I go backwards, till I see the Top of the Pole in the Middle of the Water, as at B. This done, I measure the Distance from my Station at B to the Bason at A, which suppose 72 Inches; and also measure from the Bason to the Bottom of the Pole at D, and find it 175 Inches. Next I measure the Height of the Eye from the Ground, which suppose 60 Inches. Then say, by the Rule of Three,

As the Distance from my Station to the Bason,

Is to the Height of the Eye,

So is the Dist. from the Bason to the Foot of the Pole,

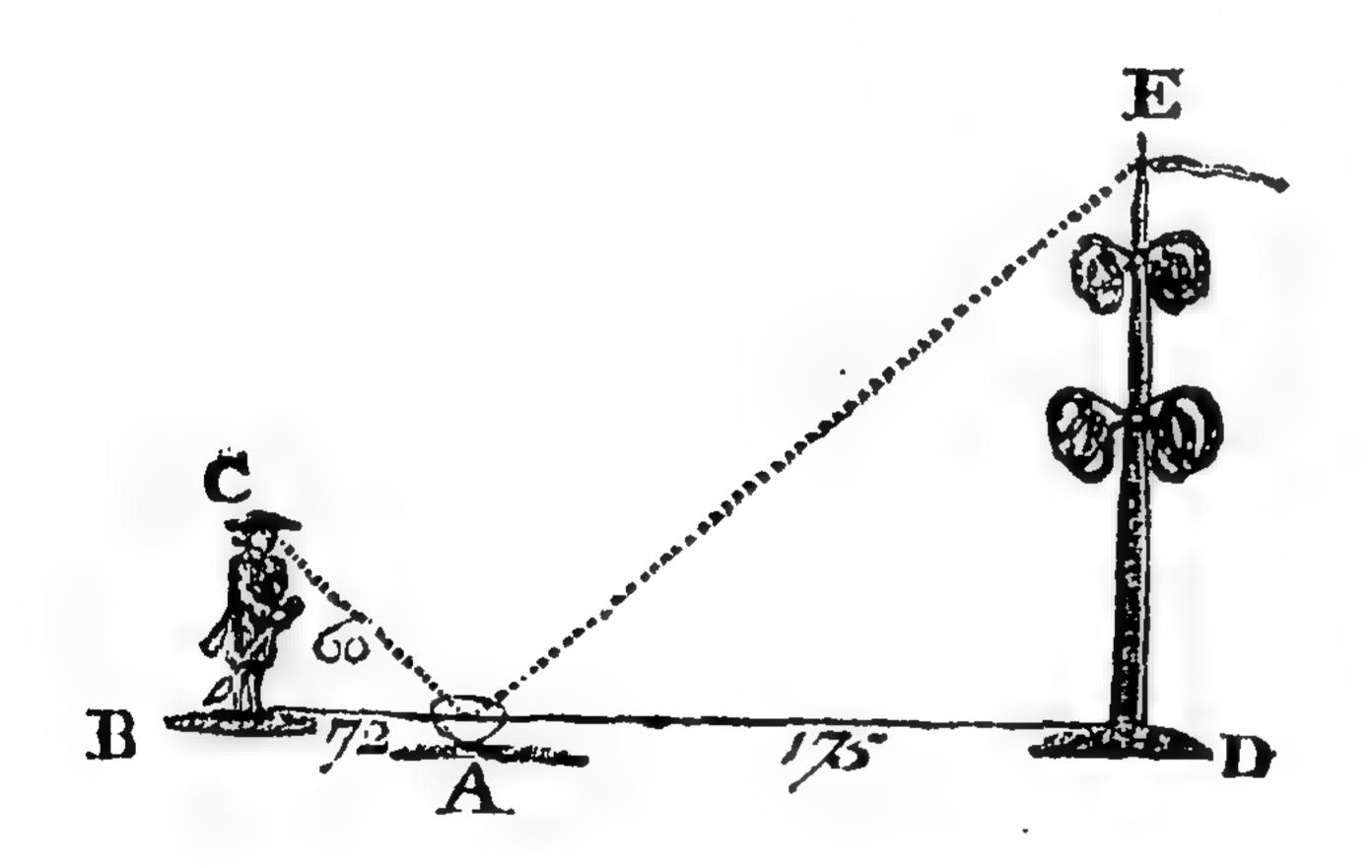
To the Height of the Pole required,

= 72

= 60

= 175

= 145.8



The same Thing may be obtain'd by a Looking Glass, laid truly Horizontal, or level on the Ground, by walking back till you can see the Top of the Building, &c. in the Middle of it, as was done by the Water.

Post Buckeye

## PROBLEM XV.

To take the Distance of the Sun, Moon, or any of the Heavenly Bodies.

Suppose it was requir'd to find the Distance of the Sun, in Diameters of his Body from us.

With a Quadrant nicely graduated, take the Altitude of the lower and upper Limb in Degrees and Minutes, and subtract the one from the other; the Remainder will give the Diameter of the Sun, which we will suppose, in this Case, to be 32 Minutes.

Then have we given in the Triangle ABC, Right Angled at B,—the Angle at A = 16', and the Side CB = .5 = Half the Diameter of the Sun, whose whole Diameter we will call 1; to find the Distance, or Side AC.

05

.26) 2865 (110 Diameters; and so far is the Sun of its own Breadths from us in the

Winter\*, but in Summer, the Angle being a little smaller, he

must, consequently, be a little surther from us.

Having found the Distance of any Heavenly Body in its own Diameters from us; you may easily tell its Distance in Miles, if you first know the Diameter of that Body in Miles. For, the Distance in Diameters, multiply'd by the Miles in one Diameter, gives the Distance fought.

For the Use of the Learner, I have here subjoined a Table of the Diameters of all the Planets in English Miles; whose Distances he may calculate at his Leisure.

Sun 800.000--- Mercury 2460--- Venus 7905--- Earth 7970---- Mars 4444 --- Jupiter 81.155--- Saturn 67.870--- Moon 2175.

In this Manner we can tell the apparent Distance of any of the Heavenly Bodies: For the Sun appearing about 1 Foot in Diameter, his apparent Distance can be only 110 Feet, or 37 Yards. The apparent Distance of the Moon is nearly the same.

I PROBLEM

## PROBLEM XVI.

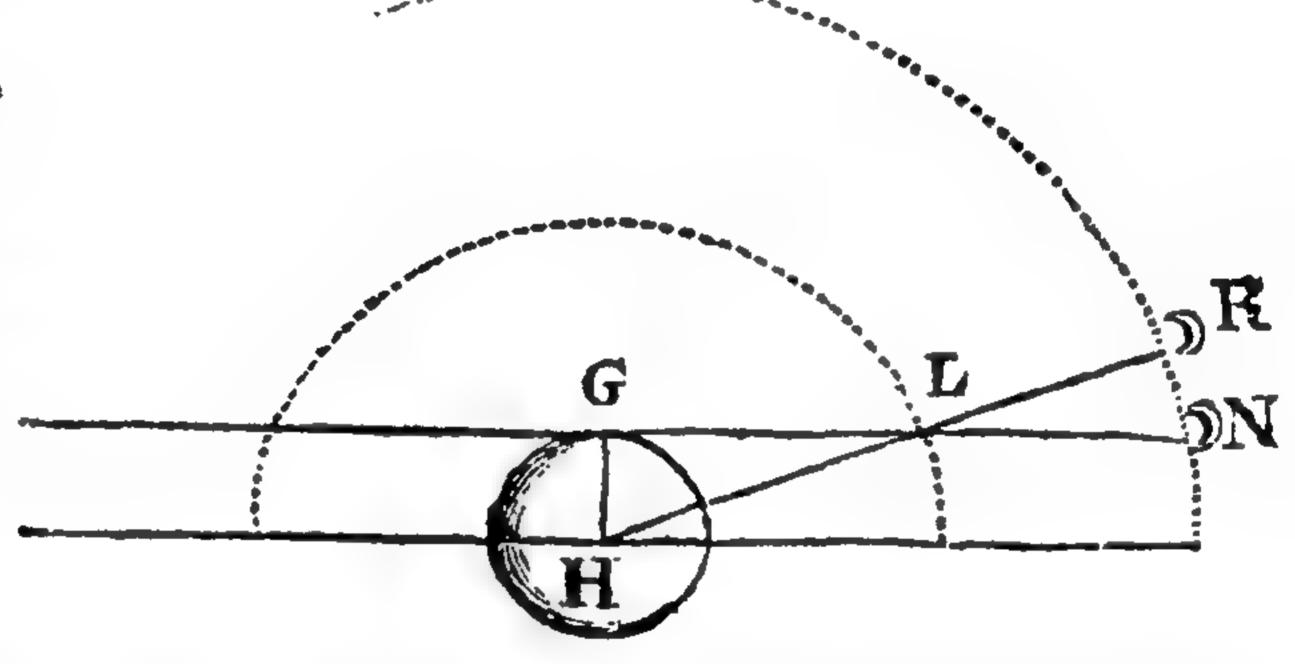
To find the Distance of the Moon from the Earth another Way.

As the Places of all the *Heavenly Bodies* are computed from the *Center* of the *Earth*, but we are obliged to view them from the *Superfices*, they must therefore appear something lower in the Heavens than they really are.

Thus, suppose, for Instance, the D to be at L in the visible Horizon; an Observer at G will see her in the Line GN, but an Eye in the Center will see her in the Line HR.—The former is her apparent Place, known by Observation with exact Instruments; the latter her true Place, and known by Calculation from Astronomical Tables of her Motion. The Difference between these two Places is the Measure of the Angle GLH, which is called the Horizontal Parallax. This Angle has been found, when the D was at a mean Distance, to be 57 Minutes nearly.

Then in the Triangle HGL, Right Angled at G, we have the Angle at L = 57', the Side GH = the Earth's Semidiameter given; to find HL, which is done thus.

Ang. L: Semid. GH:: N. Rad.
As .95 --- 1 --- 57.3



•95)57-30(60 Semidiameters; which multiply'd by the Earth's Semidiameter = 4000 nearly, you have 240000 Miles the Distance of the Moon sought.

When the D is at her least or greatest Distance from the Earth, she will be about 3 Semidiameters of the Earth nearer or farther off, than at her mean Distance.

In like Manner may the Distance of any of the Planets, Comets, or other cælestial Phænomenon be determined, by obtaining its Parallax in the Horizon.

NOTE. To find the Distance of any Place to which the Sun, Moon, or any Star is vertical, i. e. over its Head, or in its Zenith. See my Geography, p. 30.

## PROBLEM XVII.

To determine the Distance of the Sun, and all the Planets, more accurately than in the last Problem.

The Horizontal Parallax of the Moon being very difficult, if not impossible, to determine with Accuracy, on Account of the Uncertainty and Mutability of the Horizontal Refractions, which are varying according to the State of the Atmosphere: Astronomers have purfued other Methods of doing it. That which seem'd to bid the fairest—was to determine the Parallax of Venus, at a Time that, that Planet TRANSITS the Sun's Disk. This Method was first proposed by Dr. Halley, and many accurate Observations were made in distant Parts of the Globe, according to his Proposal. The Result of the several Observations made here and abroad was—that the Parallax of the Sun on the Bay of the Transit, June 6, 1761, was 8 .52. At which Time the Sun was nearly at his greatest Distance from the Earth: Consequently, has Parallax at his mean Distance will be something more, viz. 8.55.

Let S represent the Sun, E
the Earth; then in the Triangle SCD Right Angled at C;
these are given the Angle at S
= 8.65, (the Angle under
which the Semidiameter of the



Earth's Semidiameter;—to find the Distance SD or SC. Thus,

Angle S: Side CD:: Nat. Rad.: Side SD

As, 0024 --- 57.3 --- 23875 Distance of the Sun in Semidiameters of the Earth,

which multiplied by the Earth's Semidiameter gives upwards of 95 Millions of Miles.—The Sun, and consequently all the Planets, by this Observation of the Transit of Venus are found to be further off, than by all former Observations, about & Part or something more.

The Table at Page 33 of the View of the Heavens, exhibits the Distances of the Planets, according to the late and present Calculations.

The Distance of the nearest Fix'd Star is so immense, that it cannot be ascertain'd by this Method.

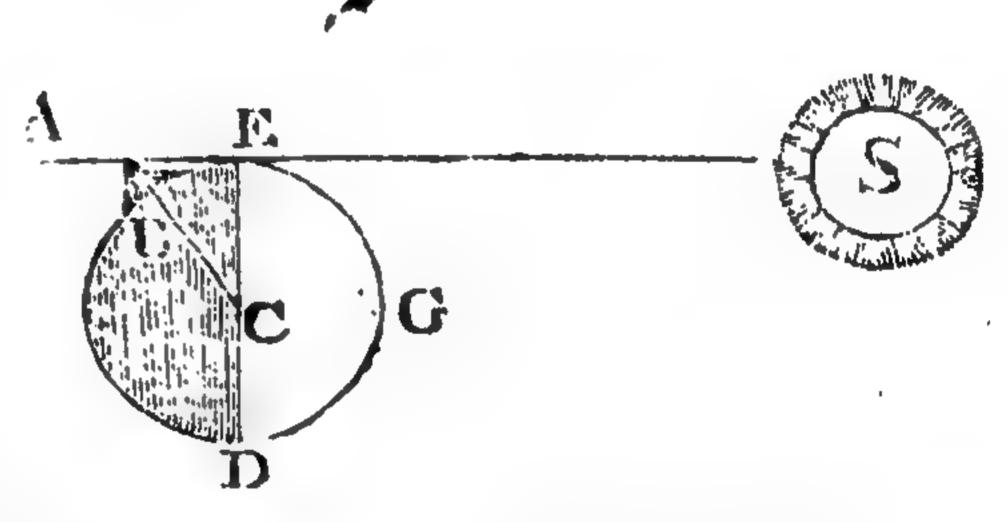
## PROBLEM XVIII.

To measure the Height of a Mountain in the Moon.

The Moon is supposed to be composed of Land and Water as our Earth is; consequently there must be some Unevennesses, or Inequalities, of Hills and Vallies as here. This indeed is confirmed by viewing her thro' a good Telescope; for then we find, that the Line, which seperates the Light from the Dark Parts on her Surface is not even or regular, but tooth'd and jagg'd with innumerable Breaks; and even in the Dark Parts, near the Borders of the lucid Surface, there are seen some small Spots enlighten'd by the Sun, which are very visible when the D is three or four Days old, and which can be nothing else but the Tops of Mountains or Rocks; since it is impossible for the Sun's Rays to fall upon those Parts only, unless they were higher than the Rest of the Surface.

The Lunar Mountains are found to be higher, in Proportion to the Body of the 1 than any Hills upon our Globe. The Manner of calculating their Heights is this.

Let EGD be the Surface of the Dand ECD the Diameter of the Circle bounding Light and Darknefs. A the Top of a Hill within the dark Part, when it first begins to be illuminated by a Ray of Light coming from the Sun at S. Then observe with a Te-



lescope the Proportion of the Right Line AE, (i. e. the Distance of the Point A from the lucid Part) to the Diameter (or Semidiameter) of the DED, for that being ascertain'd, you have in the Triangle AEC, Right Angled at E, (where the Ray of Light touches the D,) the two Sides AE and CE, to find the Hypothenuse AC, from which subtracting BC = EC, there will remain AB the Height of the Mountain.

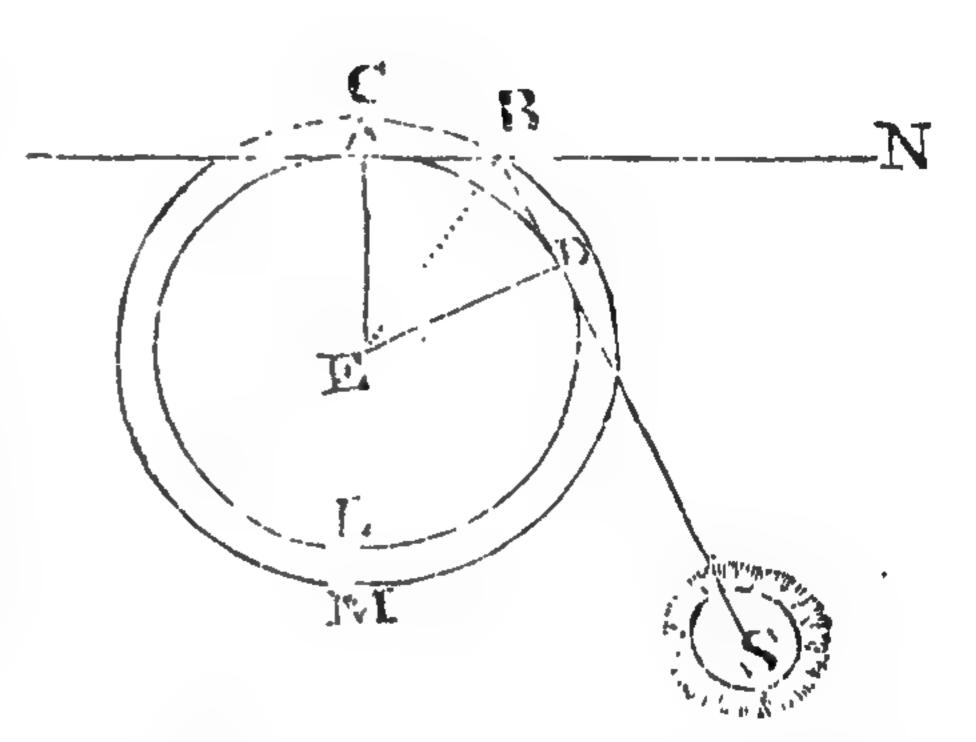
Ricciolus, on viewing the D when about four Days old, observ'd the Top of a Hill called Saint Catherine, near the N. Part of Mount Taurus, (see my Astronomy) to be illuminated, and that it was then distant from the Surface about to of the Moon's Semidiameter. Now as the Semidiameter of the DEC is about 1088 Miles, the Line AE being to of it, must be = 136 Miles. Consequently, if the of EC and of AE be added together, and then the Root of it be extracted, it will give the Line AC, from which subtracting the Moon's Semidiameter BC or CE, the Remainder, which is 8 Miles will be the Height of the Mountain sought.

## PROBLEM XIX.

To measure the Height of the Atmosphere.

The Atmosphere is that Circle of vaporous Air surrounding the Earth, which being illuminated by the Sun's Rays makes the Brightness and Glory of the Firmament we behold, whilst the Sun continues above the Horizon. And, after the Sun is gone down, the Atmosphere, which is higher than we are, will still continue to be illuminated by those Rays passing by the Earth over our Heads; but this Brightness grows less and less as the Sun descends lower, till he arrives at 18° below the Horizon; when all the Parts of the Air above sall out of his Rays, and, consequently, become dark.

To make this plainer; suppose the inner Circle ADL represents the Earth, and the Circle CBM the Atmosphere. Suppose a Person standing upon the Earth at A whose fensible Horizon is AN. Also let SB be a Ray of Light coming from the Sun, touching the Earth at D, which falls upon the distant Part of the Air, in the Horizon at B, at which time Twilight ceases. This has been found to happen when the Sun is designed.



cended 18° below the western Horizon in the Evening; and also when he is approached within 18° of the eastern Horizon in the Morning.

Now as the Arch AD is 18°, we have, by drawing the Line EB, two equal Triangles, from either of which we may find the Height of the Atmosphere required.—For in the Triangle ABE, Right Angled at A, we have given the Side AE the Earth's Semidiameter, and the Angle AEB =  $9^\circ$  = half the Arch of the Sun's Descent below the Horizon, and the Angle ABE =  $81^\circ$ , to find the Hypothenuse EB, from which if you subtract the Semidiameter of the Earth, the Remainder will be the Height of the Atmosphere.

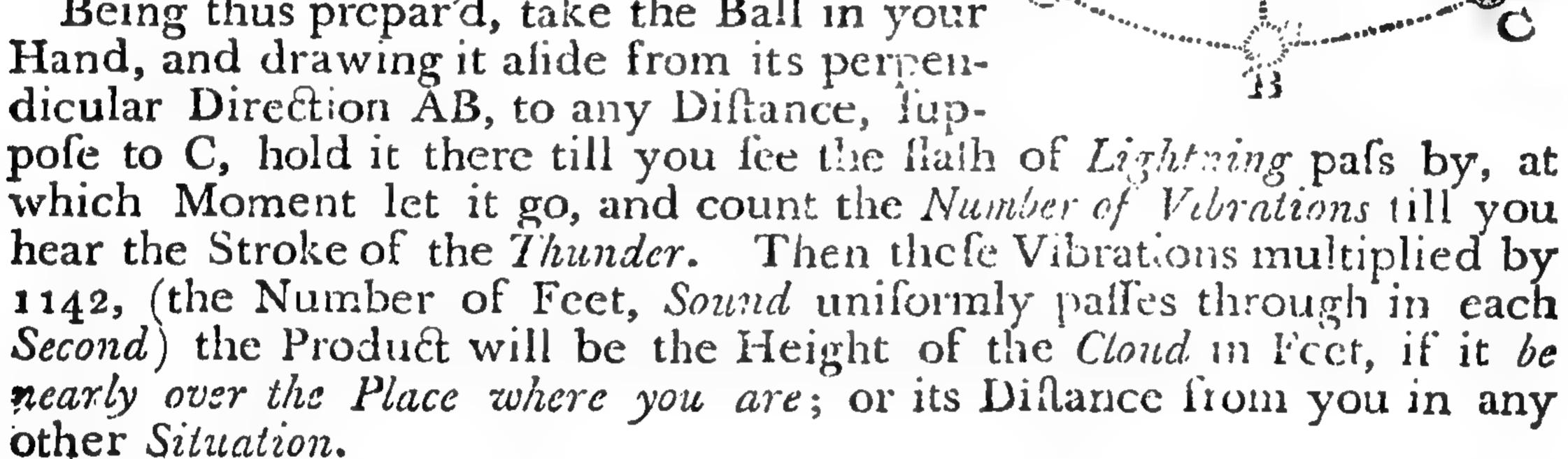
## PROBLEM XX.

To measure the Distance of a Cloud, from which issues Lightnings and

Take a small Ball of Lead, Ivory, or any other matter, and affix it to the End of a fine Thread. Then measure from the Center of the Ball, along the Thread, exactly 39.2 Inches, where make a Loop. This done,

fuspend it by that Loop to the Ceiling of the Room, or to any other Place where it may hang freely, and vibrate backwards and forwards like a Pendulum, as in this Figure. Now the Property of this little Instrument is, that each Vibration, whether it passes through a larger or finaller Space, will be performed in one Second of Time.

Being thus prepar'd, take the Ball in your



Thus, suppose the String is found to make 8 Vibrations between the Lightning and the Thunder; then 8 x 1142 = 9136 Feet, which, divided by 5280 (the Feet in 1 Mile) gives 14 Mile nearly; and so far is that alarming Tempest from you.

In this Manner you may continue to measure the Distance of the Cloud all the Time it passes from your Zenith to the Horizon, and by that Means be acquainted with the Danger it seems to threaten the Neighbourhood, as well as the Extent of the visible Hemisphere of Clouds.

The Distance also of a Ship at Sea, or a Fort, may be estimated in the same Manner, by counting the Vibrations from the Flash of the Powder to the Report of the Gun.

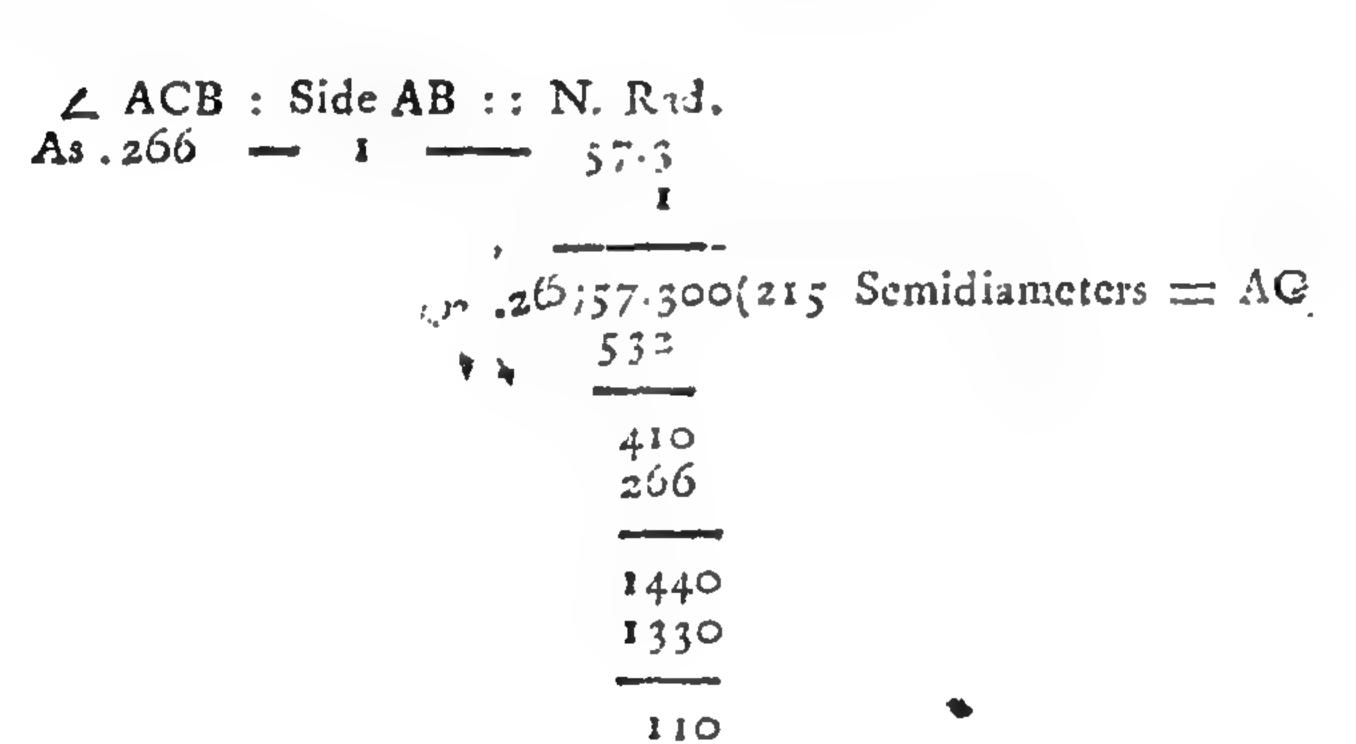
PROBLEM

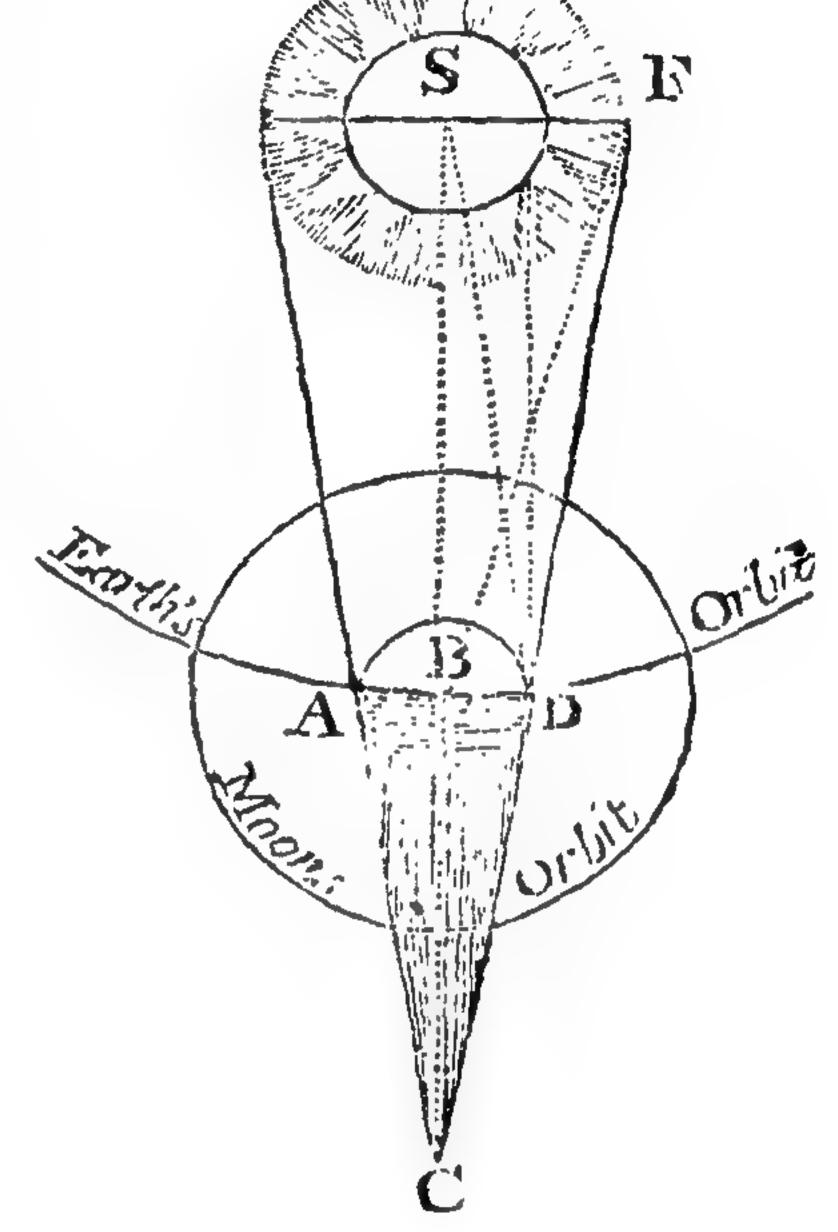
## PROBLEM XXI.

To calculate the Length of the Earth's or Moon's Shadow.

The Angle of the Cone ACD of the Earth's Shadow in the annex'd

Figure, is equal to the Sun's apparent Diameter\*, which, at a mean Distance from us, is about 32 Minutes. Hence, in the Triangle ACB, Right Angled at B, we have the Angle ACB = 16 Minutes, the apparent Semidiameter of the Sun, and AB the Semidiameter of the Earth = 1; to find AC or BC the Length of the Shadow, which is done thus.





Thus, when the Sun is at a mean Dislance from us, the Shadow of the Earth reaches about 215 Semidiameters beyond it: But when the Sun is at his greatest or least Distance, the Shadow will be lengthened or shortened 3 or 4 Semidiameters, more or less.

Hence, you may also determine the Height of the Moon's Shadow: For, as the Moon is never at any great Distance from the Earth, the apparent Semidiameter of the Sun must be nearly the same there as here. Consequently, the Moon's Shadow must contain the same Number of Semidiameters of the Moon, as the Earth's Shadow does Semidiameters of the Earth: Which Semidiameters, multiply'd by the Miles in the Semidiameter of the Moon or Earth, will give the Length of the Shadow respectively in Miles.

The Semiangle of the Cone of the Earth's Shadow BCD, is equal to the apparent Semidiameter of the Sun view'd from the Top of the Shadow, which Angle is always equal (in the Shadow of every Planet) to the apparent Semidiameter of the Sun SBF, lessen'd by his Horizontal Parallax BSD at that Planet. But as the Horizontal Parallax of the Sun, i. s. the Angle under which the Earth is seen from thence, is scarcely 10 Seconds, it may be omitted, as is done in the above Calculation.

## ROBLEM

To calculate the Diameter of the Earth's Shadow at the Distance of the Moon; and also, the Diameter of the Moon's Shadow at the Earth.

In the following Figure let S represent the Sun E, the Center of the

Earth, M the Moon, EC the Cone of the Earth's Shadow (at a mean) = 215 Semidiameters of the Earth: Then MC will be the Cone of the Earth's Shadow reaching beyond the Moon, whose Length is thus found.

Semidrs. From EC the Cone of the Earth's Shadow Subtract EM the Dist. of the Moon in the Earth's Semid. Remains MC the Shadow of the Earth beyond the Moon

Then, by Reason of similar Triangles, it will always hold;

As the Length of the whole Shadow Is to the Diameter of the Earth So is the Length of the Shadow beyond the Moon "To the Diameter of the Shadow at the Moon

MC EC 39820 39820 7964

> Miles = cd, the Diameter of the Earth's Shadeev at the Distance of the Alvon. 1075

= 155

EC

ab

cd

MC

By this Problem also may be found the Diameter of the Moon's Shadow at the Surface of the Earth, and, consequently, how much of the Earth is involv'd in that Shadow in an Eclipse of the Sun. For the Length of the Moon's Shadow is found to be about 60 Semidiameters of the Earth, which is nearly the

Moon's mean Distance from us; her Shadow, in that State, must, there-'fore, reach as far as the Center of the Earth.—But as the Moon is, sometimes, almost 4 Semidiameters of the Earth nearer, the Shadow must reach 4 Semidiameters beyond the Center of the Earth; and when the Moon (as she sometimes is) is 4 Semidiameters further from us, the Shadow will then not reach the Earth at all. In fuch Case, the Sun, though centrally eclips'd, will not be totally covered by the Moon; but an Annulus, or Ring of Light, will appear round the Border of that Luminary, as happen'd April 1st, 1764.

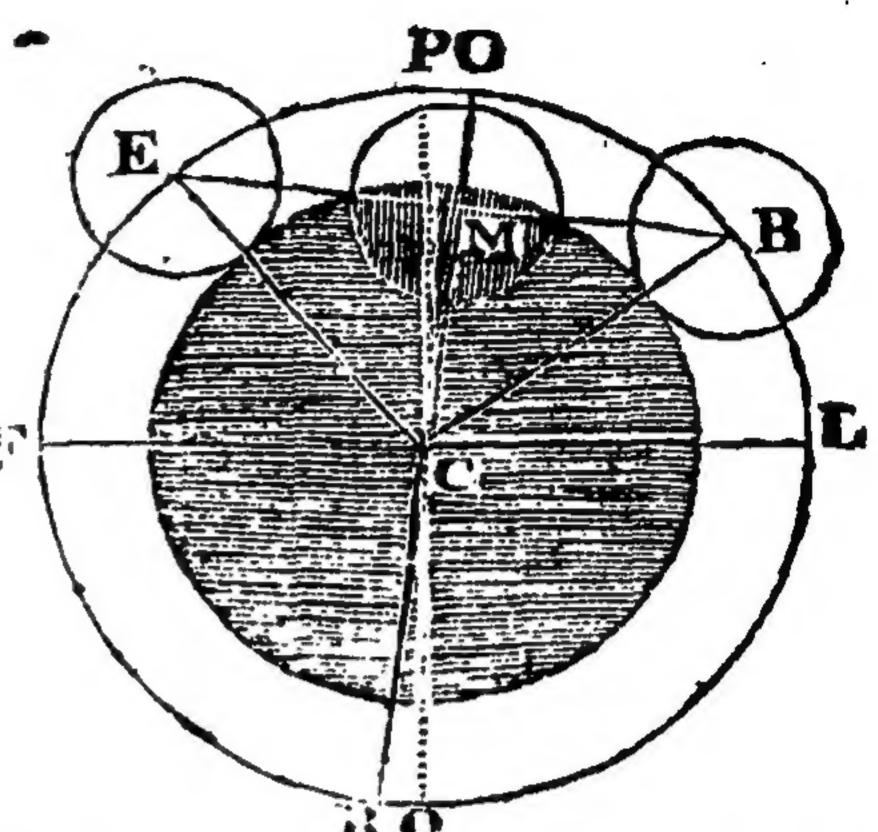
The Method of finding the Distance of the Moon, in Semidiameters of the Earth, is shewiff at Problem XVI.

## PROBLEM XXIII.

To calulate the Beginning, End, and total Duration of an Eclipse.

Having, from Astronomical Tables, obtain'd the Time of the Middle of the Eclipse, which suppose to be December 21 Day, 11 Hours, 49 Minutes, with the Latitude of the Moon at that Time, = 40 Minutes, you may then proceed to find the Beginning, End, and total Duration, as follows.

From a Scale of equal Parts, of any Size, take off the Semidiameter of the Earth's Shadow, which, at the Distance of the D (at a Mean) is about 42'; and, setting one Foot in C, describe the inner shaded Circle, to express the set of the Cone of the Earth's Shadow off at that Place where the D passes through in that Eclipse.—With the Sum of the Semidiameter of the D = 16', and Earth's Shadow = 42', (which together = 58') taken from the same equal Parts describe the outer



Circle.—Draw the Line FL through the Center, to represent the Ecliptic, or Path of the Earth's Shadow; cross it, at Right Angles, with the dotted Line PQ, to express the Poles of the Ecliptic.—Then with a Line of Chords, or a Protractor, set off 5.000 from P, upon the outer Circle, towards the Right Hand, because the Latitude of the pois North ascending, to express the Angle of the Moon's Path with the Ecliptic, and draw the Line OR.—Take the Latitude of the point the same Scale of equal Parts, and set it from C to M upon the Line CO.—Then draw a Line through M, at Right Angles to CO, and that Line will represent the Path of the polynomia the Eclipse.—Next, with the Semidiameter of the points B, M, and E, severally, the three little Circles; so will the Circle at B represent the path he Eclipse.

Now, from the Center C, draw two Lines to B and E; then in the Right Angled Triangle CMB, Right Angled at M, we have given CM, the Latitude of the D = 40, and CB = CE the Sum of the Semidiameters of the Moon and Earth's Shadow = 58; to find MB = ME, the Motion of Half Duration of the Eclipse.

From Square of 58 = 3364
Take Square of 40 = 1600

Extract the Root 1764(42 Minutes, = the Motion of Half the Duration. And because the Moon passes over 31 of these Minutes (at a mean Rate) in one Hour, we have this Proportion.—If 31 Min.: 1 Hour: 42 Min.: 1 Hour 21 Min. which subtracted from, and added to, the Middle, will give the Beginning and End.

Middle of the Eclipse Half Duration subtract and add Dec. 21 11 49

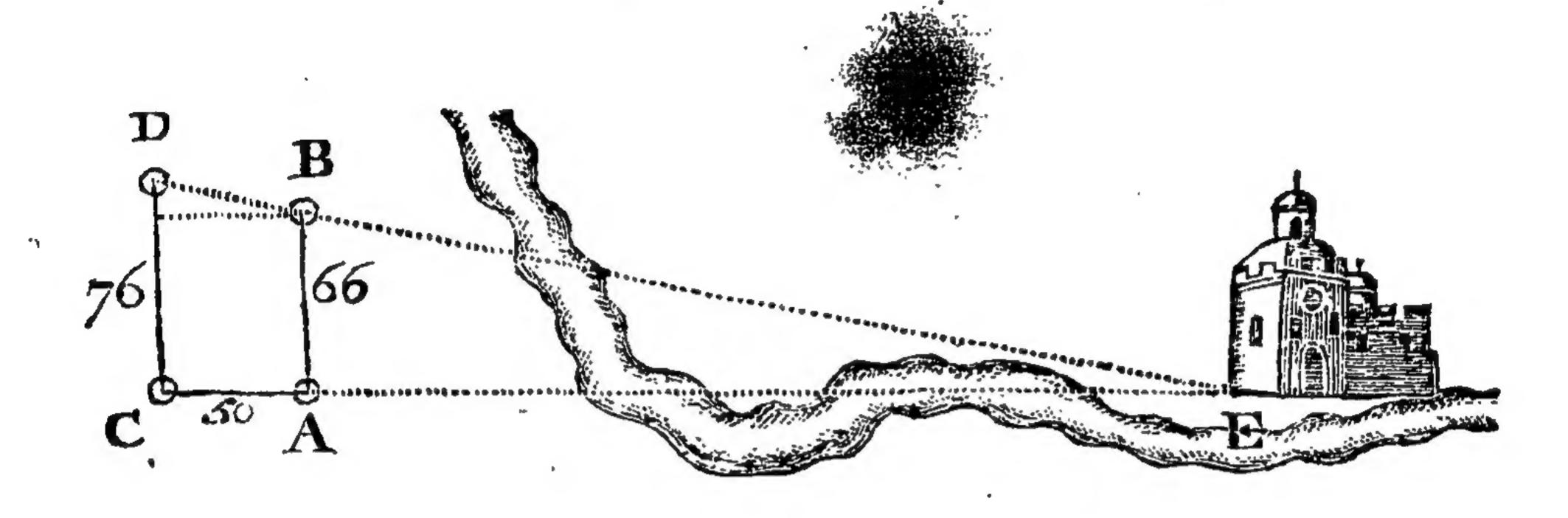
Beginning 21 10 28 End 21 13 10

Note. The Middle of an Eclipse, with the Latitude of the Moon, may be casily had from my Astronomy.

## PROBLEM XXIV.

To take the Distance of an inaccessible Object without the Help of any Instrument.

Suppose E, in the following Figure, to be a Fort, whose Distance you want to know, and you cannot approach it, on Account of some. Moat, Ditch, or River, lying between you and the Object.



First, at some Distance from the Ditch or River, set up a Stick, as at C; then advance, in a Right Line, towards E, any Number of Yards, suppose 50, and set up another Stick at A; next, move, in a Line perpendicular to CE, from A to B, any Distance, suppose 66 Yards, and set up another Stick at B; then return back to C, where you began, and remove from thence, in a Line perpendicular to CE, till you see the Stick at B and the Object E in a Right Line, and set up another Stick in that Place at D, measuring the Distance from C to D, which suppose 76 Yards. Then it will always hold;—

Note. If, in the third Term, you had us'd the Distance AB = 66 Yards, you would have obtain'd the Distance from A to E = 330.

## PROBLEM XXV:

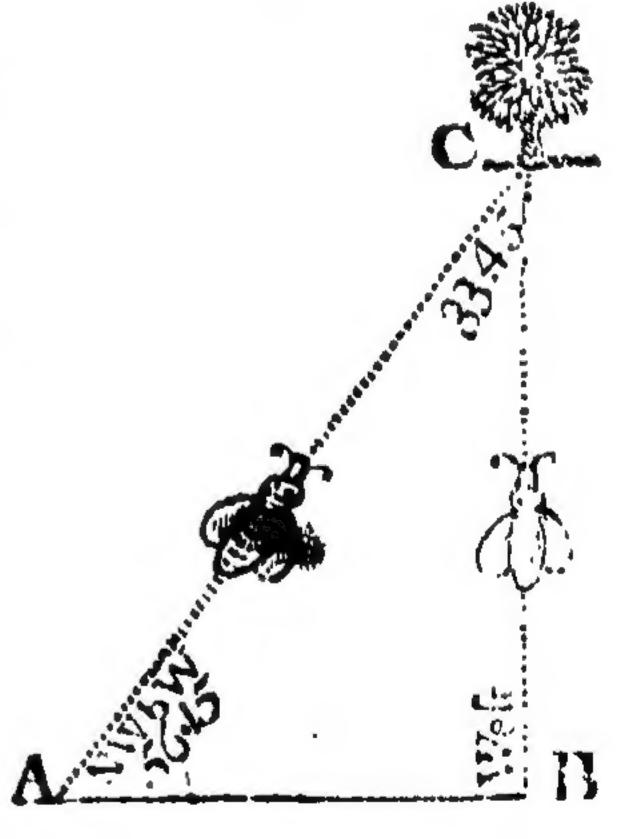
To find, by a new Method, where the Bees hive in large and extensive Woods, in Order to obtain their Honey.

Take a Plate or small Piece of Board, on which is spread a little Honey or Treacle, and set it down on a Rock or Stump of a Tree within the Wood. This the Bees will soo, find out if any are near, for it is generally believ'd they smell Things of that Nature at the Distance of a Mile, or surther. Whilst the Creatures are feeding, secure two or three of them in a Beautiful one of them go, observing care by a Pocket Compass, the Course he takes; for Bees, after they the Air, sly directly in a strait Line to the Tree where their Hive is.

Suppose, for Example, the first Bee is found to fly directly West; then you may be sure the Tree is some where in that Line from your present Station. But, in order to know how far, you must make an Offset, either North or South, as large as you can, which in this case we will suppose to be 100 Rods or Perches (the larger the better) to the South. Here you must let go another Bee, observing his Course as before, (for this Bee, being loaded like the other, will sly directly to the Hive) which Course we will suppose to be N. W. by W.—56° 15' towards the West, it only remains now to find where these two Courses or Lines intersect or meet with each other, for there you will find the Tree in which the Honey is.

This may be easily done.—For in the Right Angled Triangle ABC, are given the Right Angle B, the Course of the first Bee, the Angle at A, the Course of the second Bee, and the Distance AB; to find BC, or AC, the Distance of the Tree from either Station.

∠C: Base:: N. Rad.: Dist. AC
As 33 75 — 100 — 60.7 — 179.8 Perches.



Formerly, they found the Honey by surprizing the Bees, and sollowing them, one after another, till they found out the Hive; since this Trigonometrical Method has been us'd, the Searchers discover that Booty in a sew Hours, which before requir'd many Days.

CONCLUSION.

## CONCLUSION.

THESE sew Problems are sufficient to point out the great Use of this Branch of Learning. The Advantages resulting from it to Society are very great;—almost infinite.—Nothing however posited in the Heavens;—nothing upon the Earth or Seas;—but its Distance and Dimensions may be ascertained by it.—It is no Wonder then, that Pythagoras, a learned Philosopher of Samos, when he had discovered that samous Proposition (47th east Book of Euclid) which is the Foundation of this Science, the in Gratitude, sacrifice an ileicatomb, i. e. 100 Oxen, to the por inspiring him with such an useful Invention, which he judge tond the Power of human Abilities to discover.

Thus by one plain Geometrical Figure, having three Sides and three Angles, and allilled by the Rule of Three, you fee what amaking Truths may be discovered. This illustrates not only the old Motto,--Tria funt omnia \*---but also proves the Truth of that in the Title-page.

Cuncla Trigonus habet, satagit quæ docla Mathesis, ille aperit clausum quicquid Olympus habet.

Which may be English'd thus:

In Heaven the latent Science lay conceal'd, Till the Triangle came, and Truth reveal'd.

<sup>\*</sup> Omnium prope Deorum Patestas triplici Sieno ostendatur; ut, Jovis trisiam Fulmen, Neptuni Tridens, Plutenis Canisstriceps; vel quod Ormia ternario continentur.